

# lecture 8. multi-messenger astrophysics

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**Advanced Astroparticle Physics**

NPAC M2

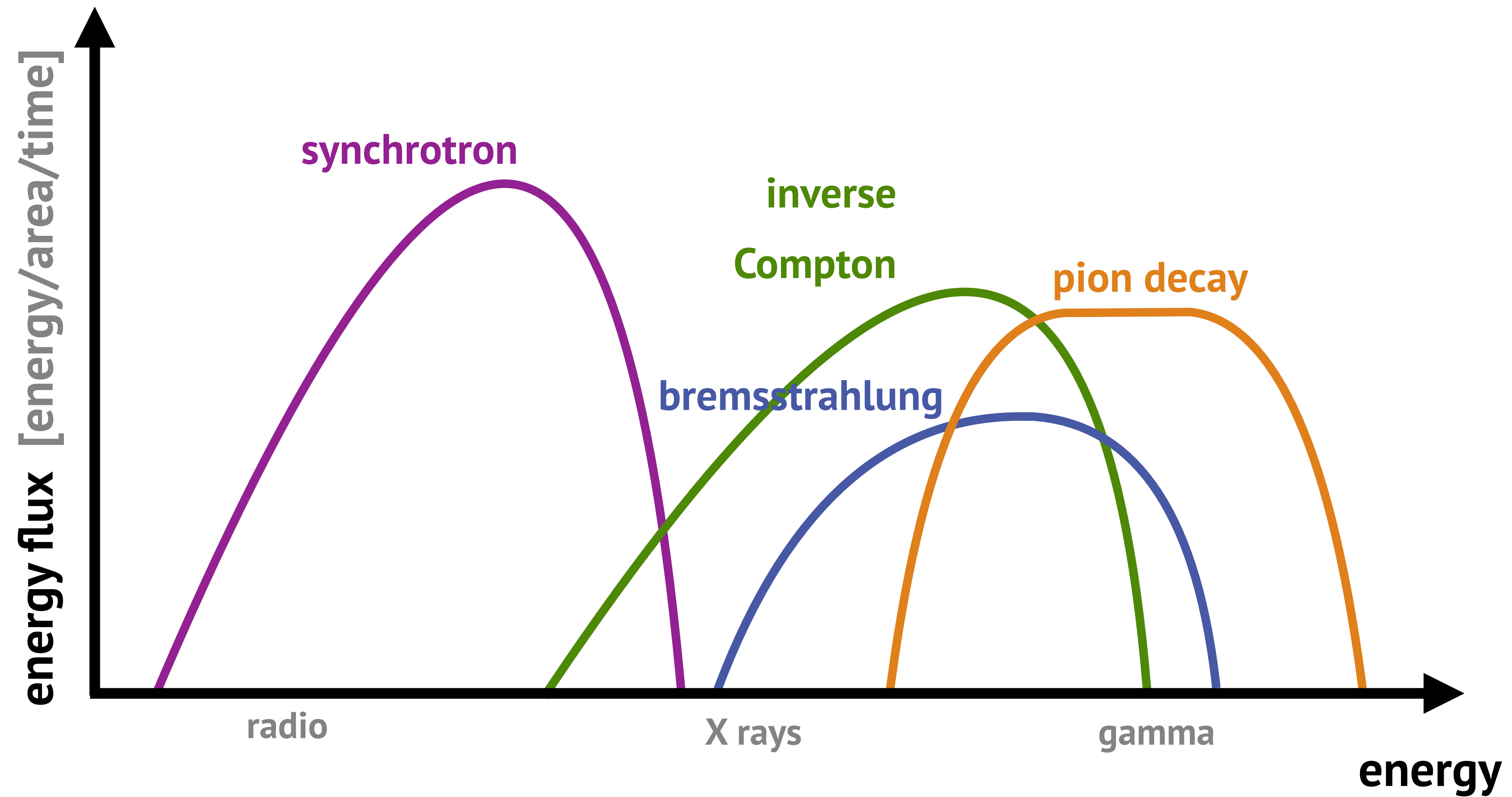
2024-2025

# in today's class...

- ▶ **interaction processes**
  - ◆ leptonic
  - ◆ hadronic
  - ◆ lepto-hadronic
- ▶ **multi-messenger observations**
  - ◆ milestones
  - ◆ networks
- ▶ **astroparticles as probes of BSM**

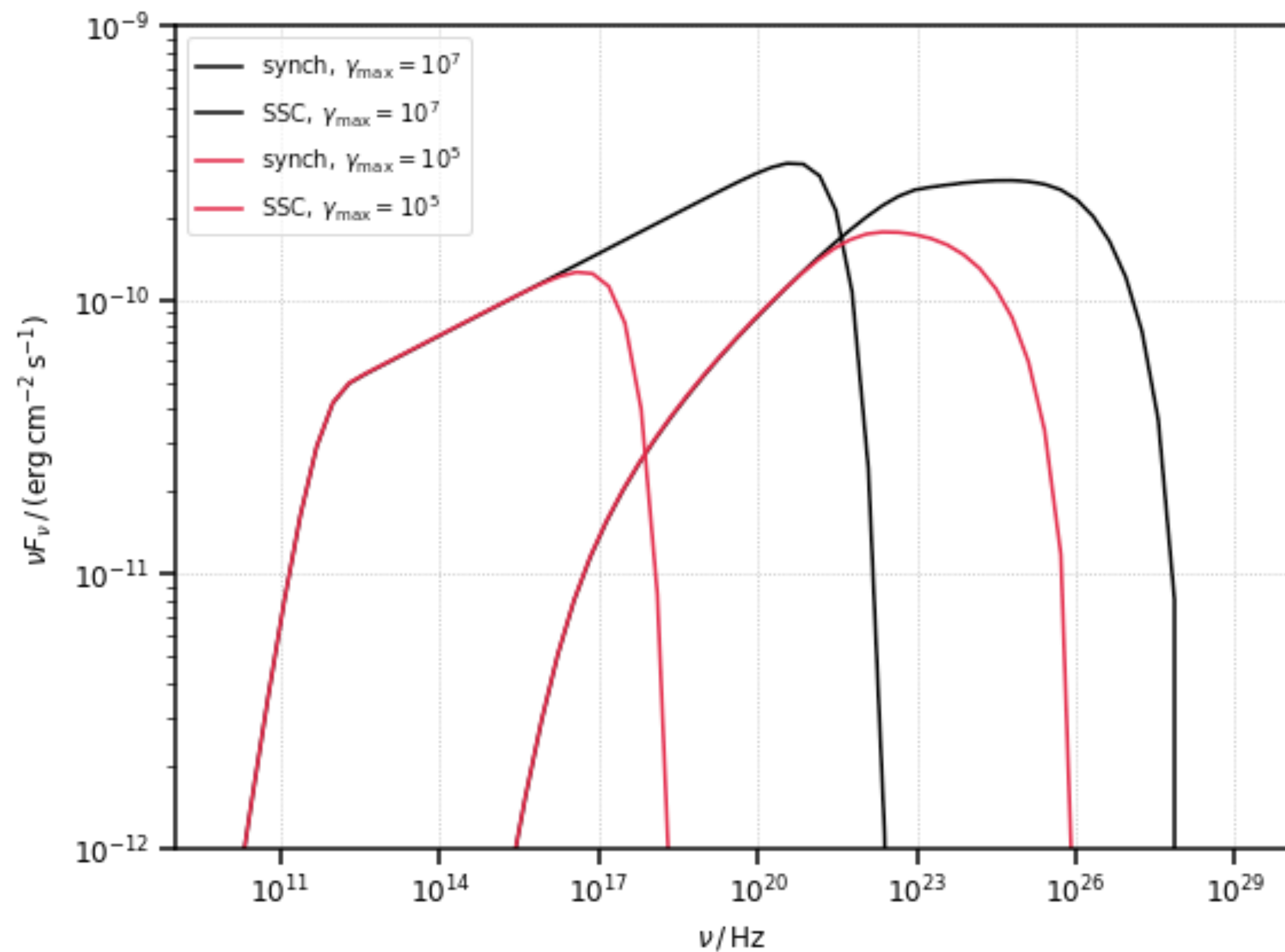
# particle interactions in astrophysical objects

# non-thermal spectral energy distribution

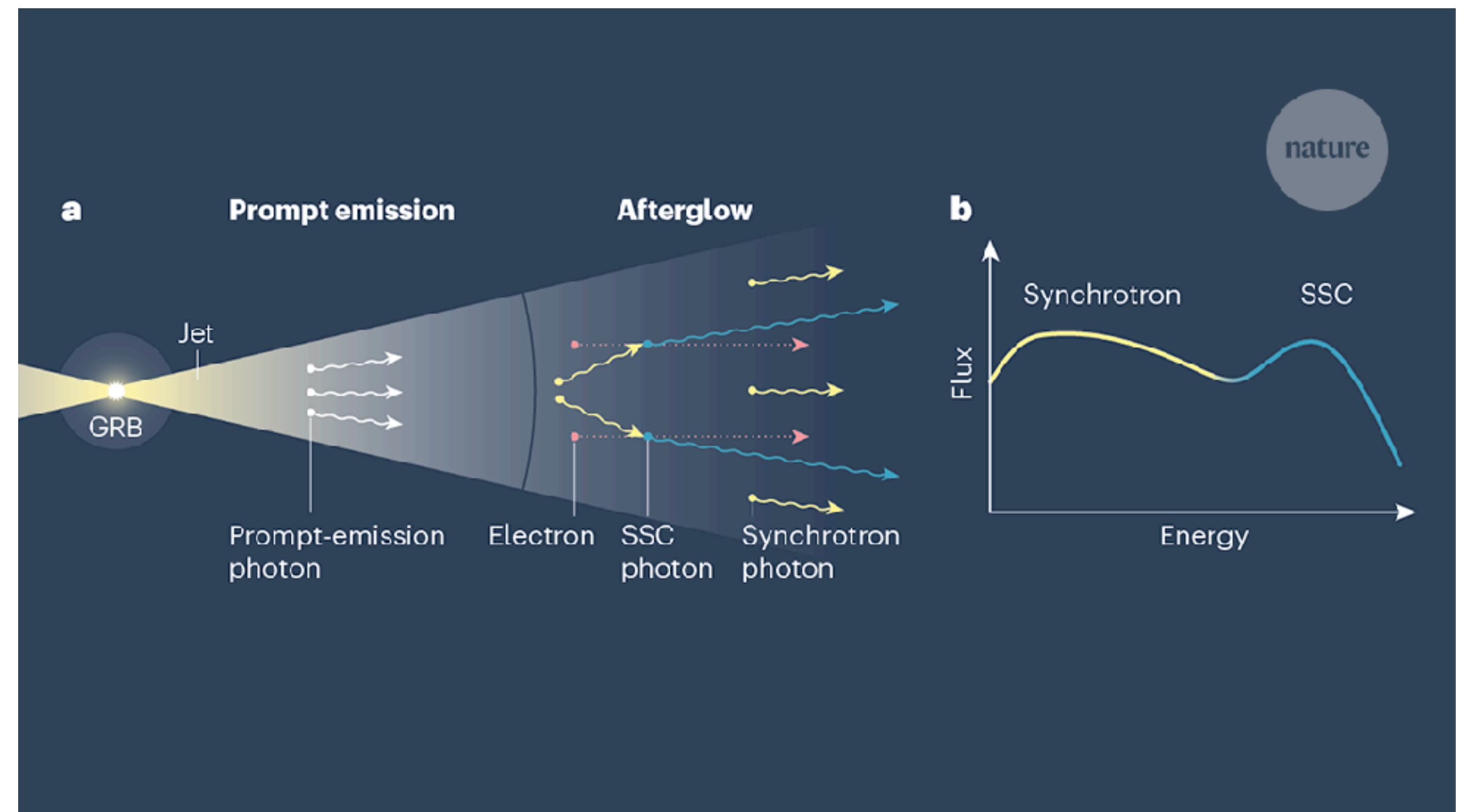


# synchrotron self Compton (SSC)

- ▶ electrons emit synchrotron radiation
- ▶ synchrotron photons become targets for Compton scattering



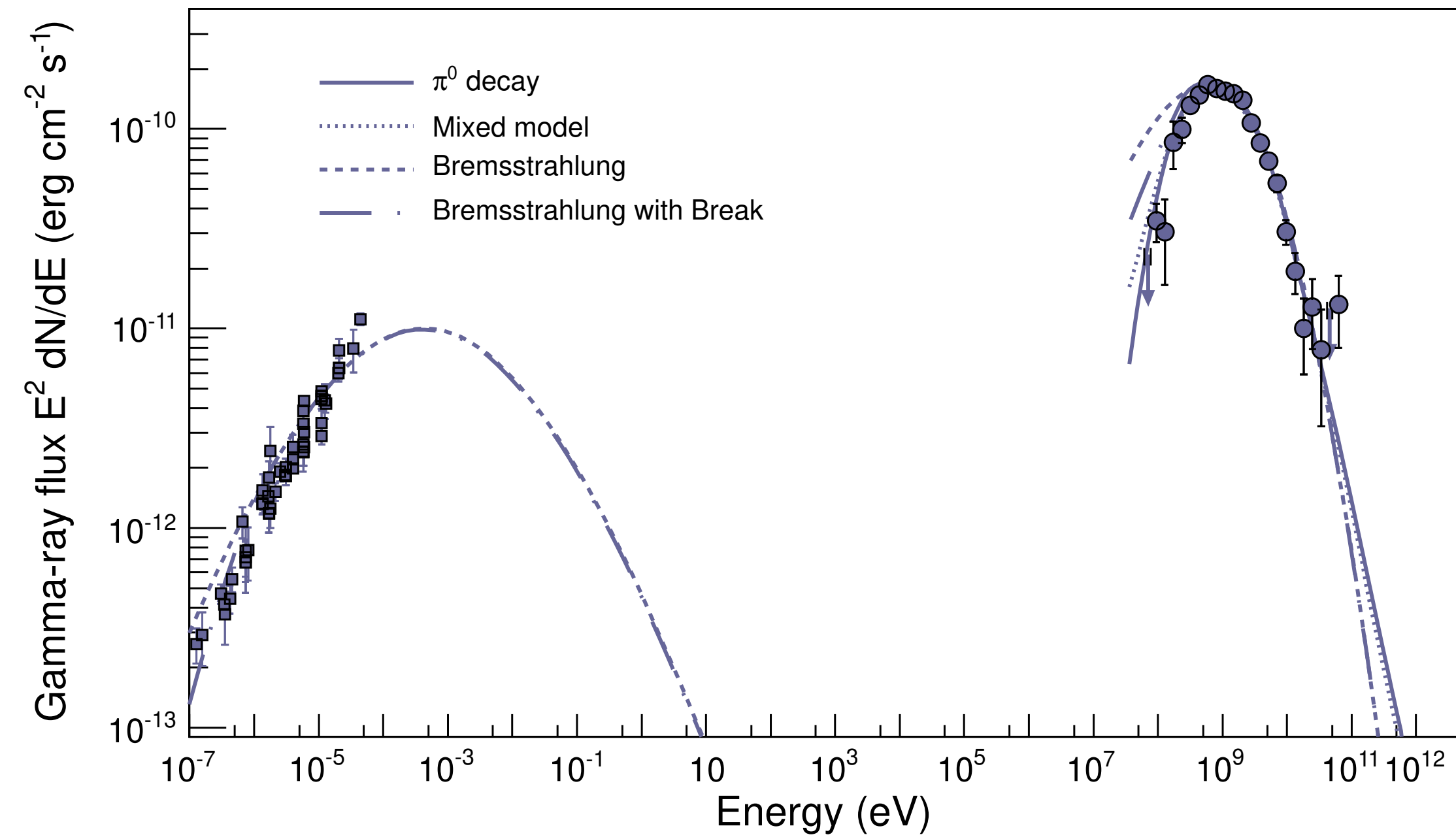
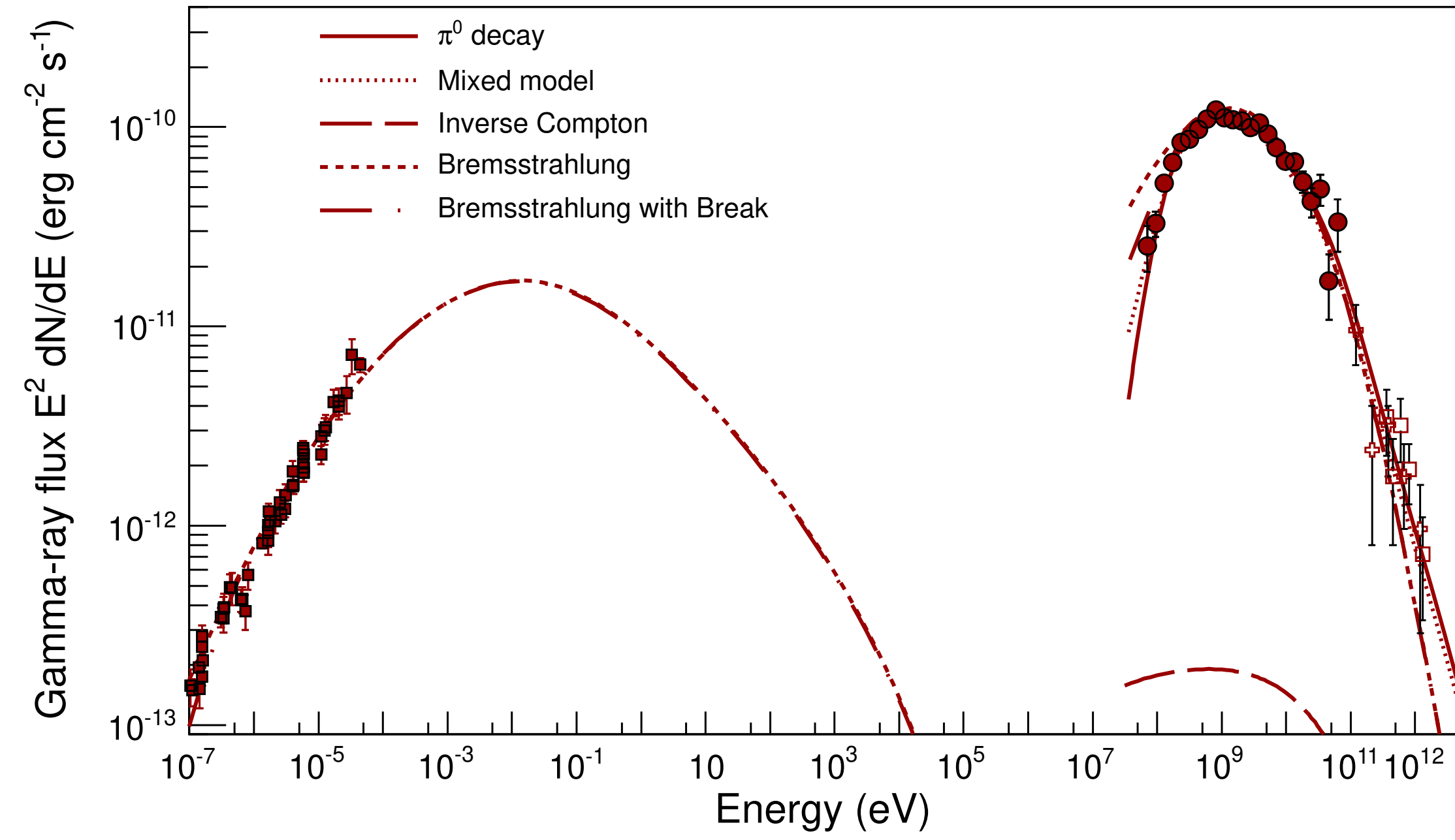
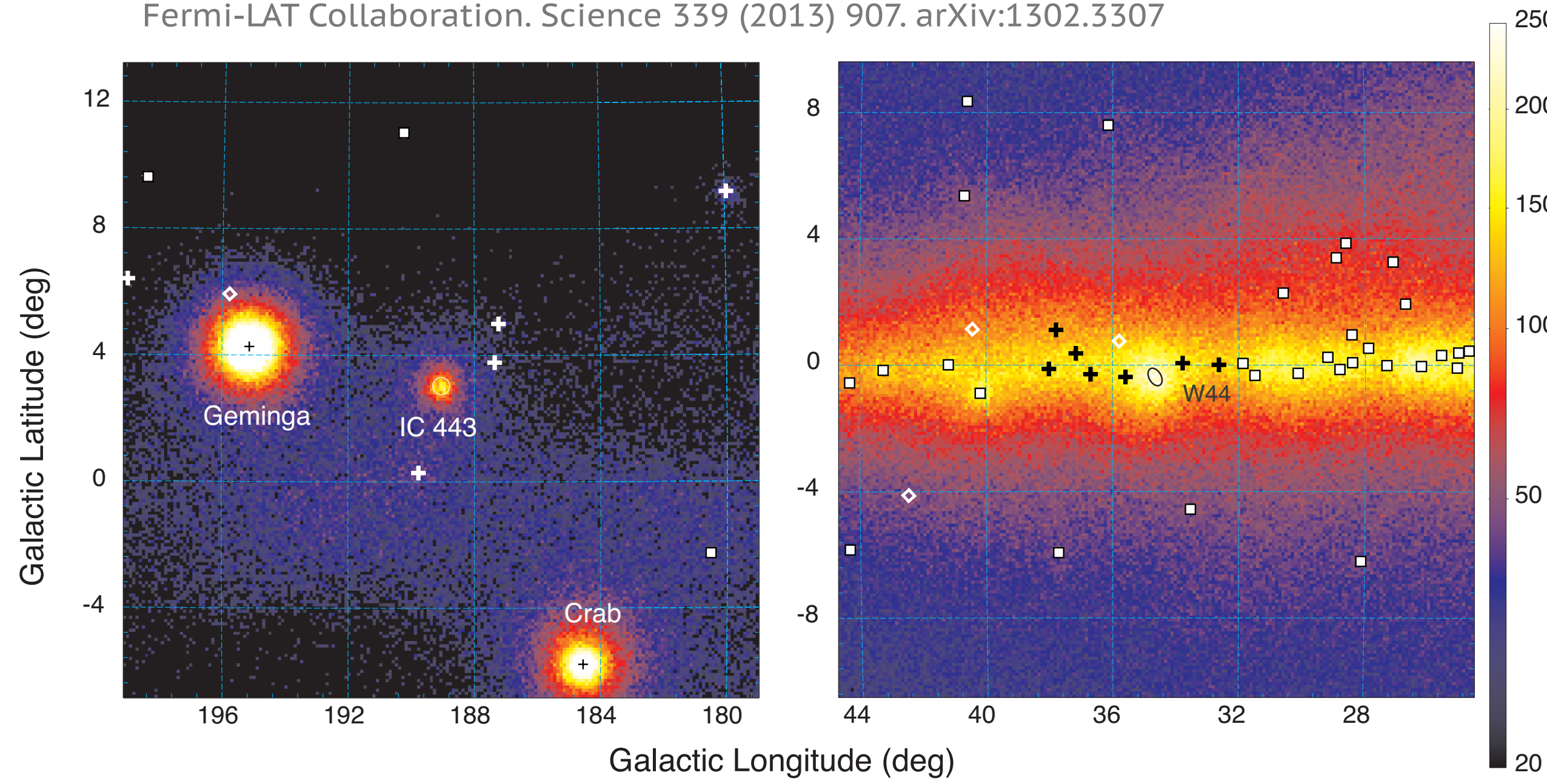
Simulated with agnpy: [https://agnpy.readthedocs.io/en/latest/tutorials/synchrotron\\_self\\_compton.html](https://agnpy.readthedocs.io/en/latest/tutorials/synchrotron_self_compton.html)



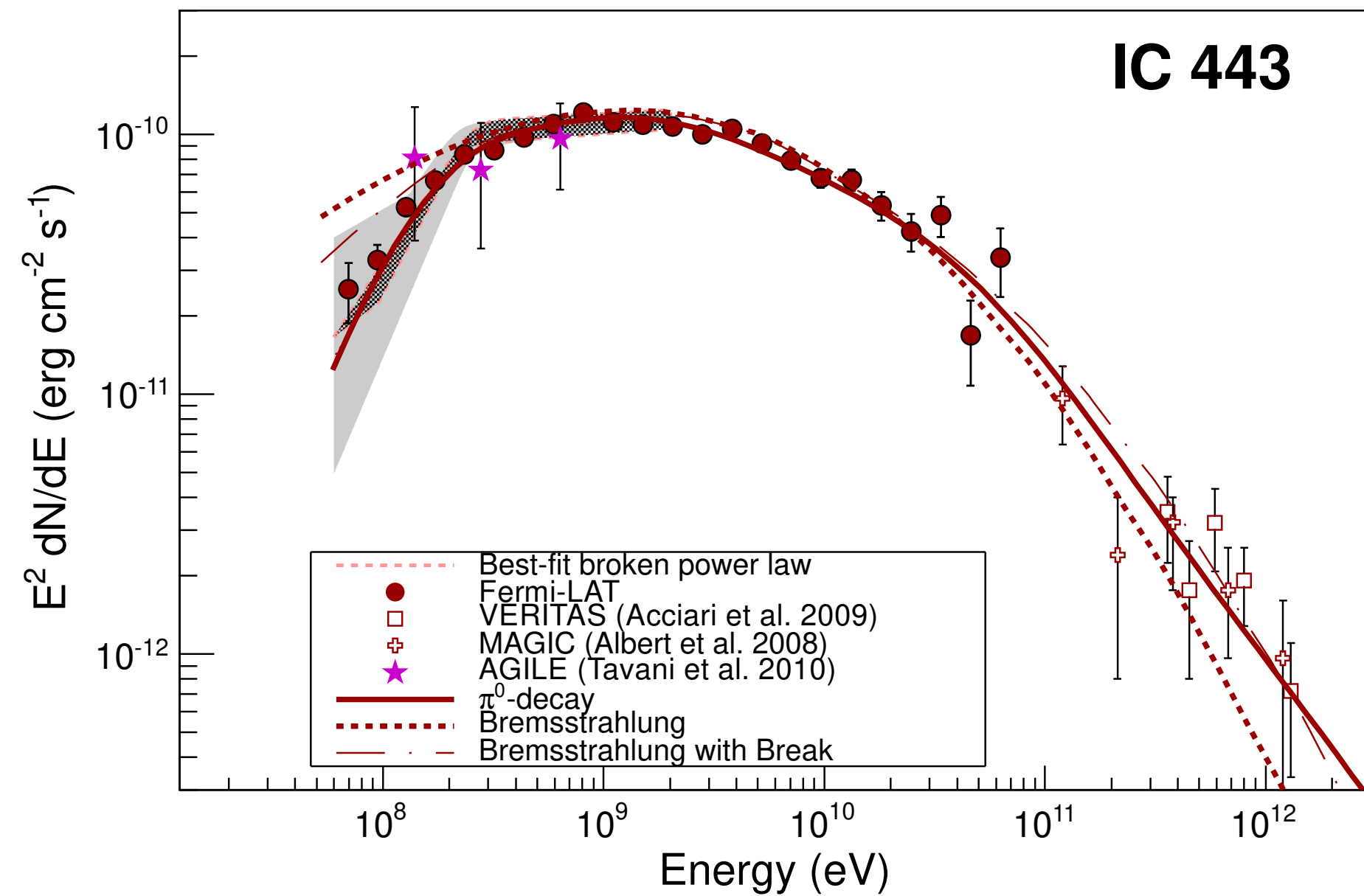
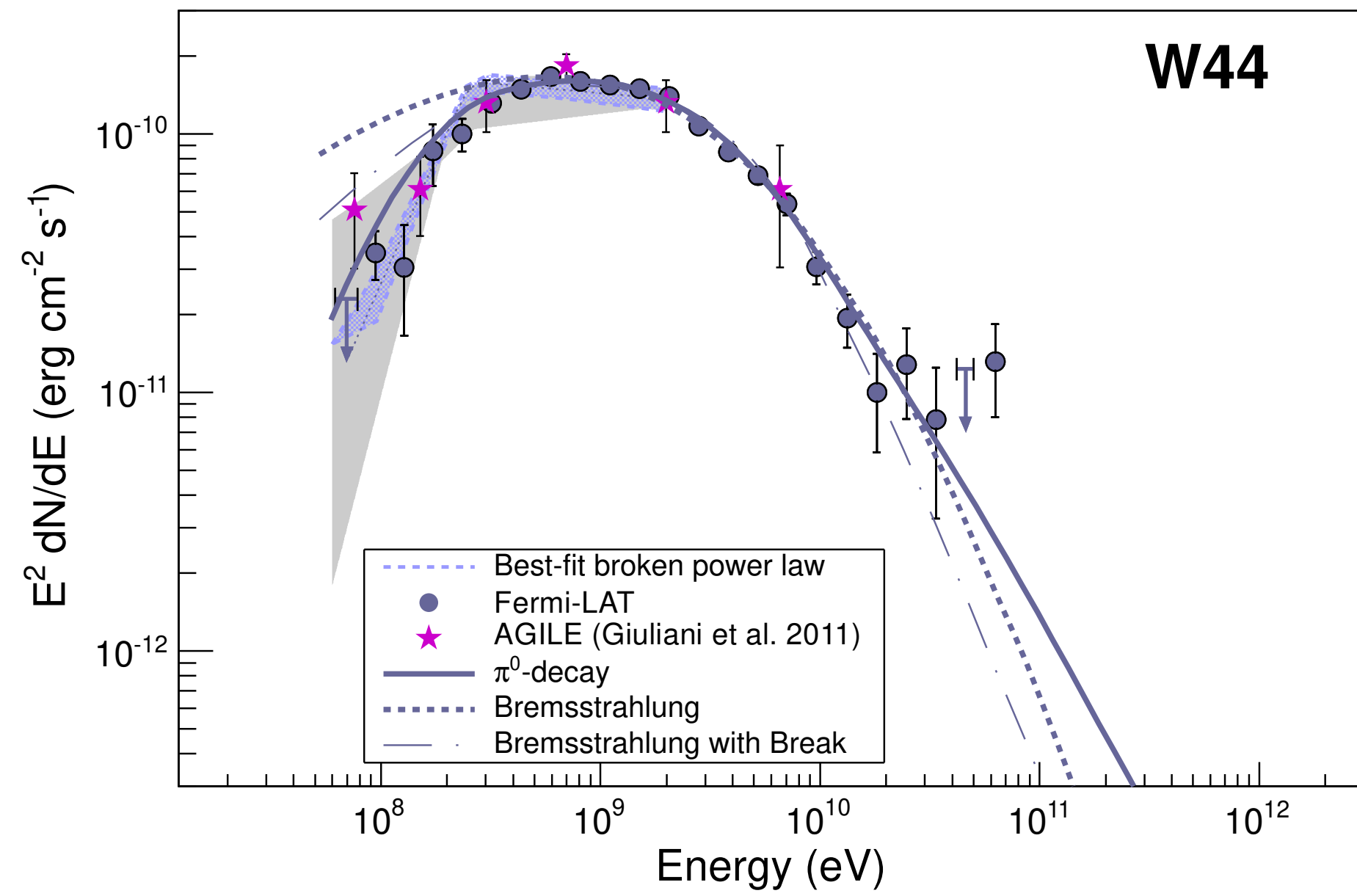
Zhang. Nature 575 (2019) 448.

# the pion bump

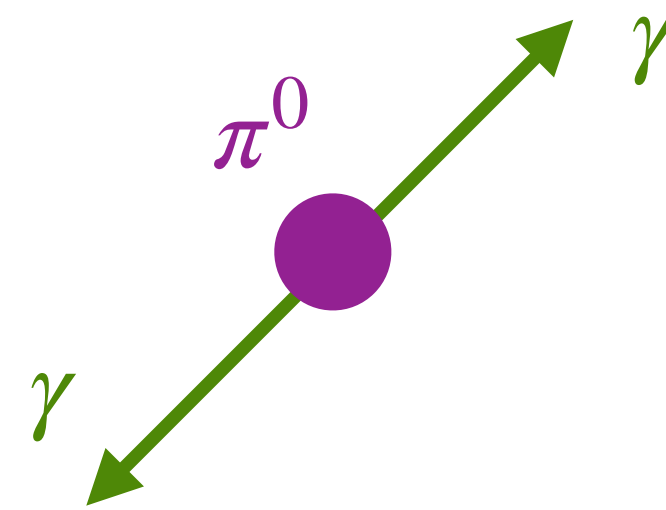
Fermi-LAT Collaboration. Science 339 (2013) 907. arXiv:1302.3307



# the pion bump

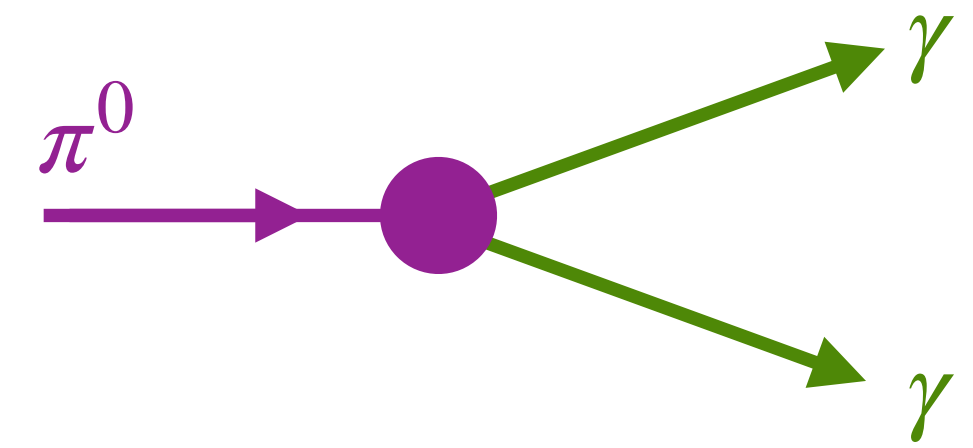


CoM frame



$$E^* = \frac{1}{2} m_{\pi^0} c^2$$

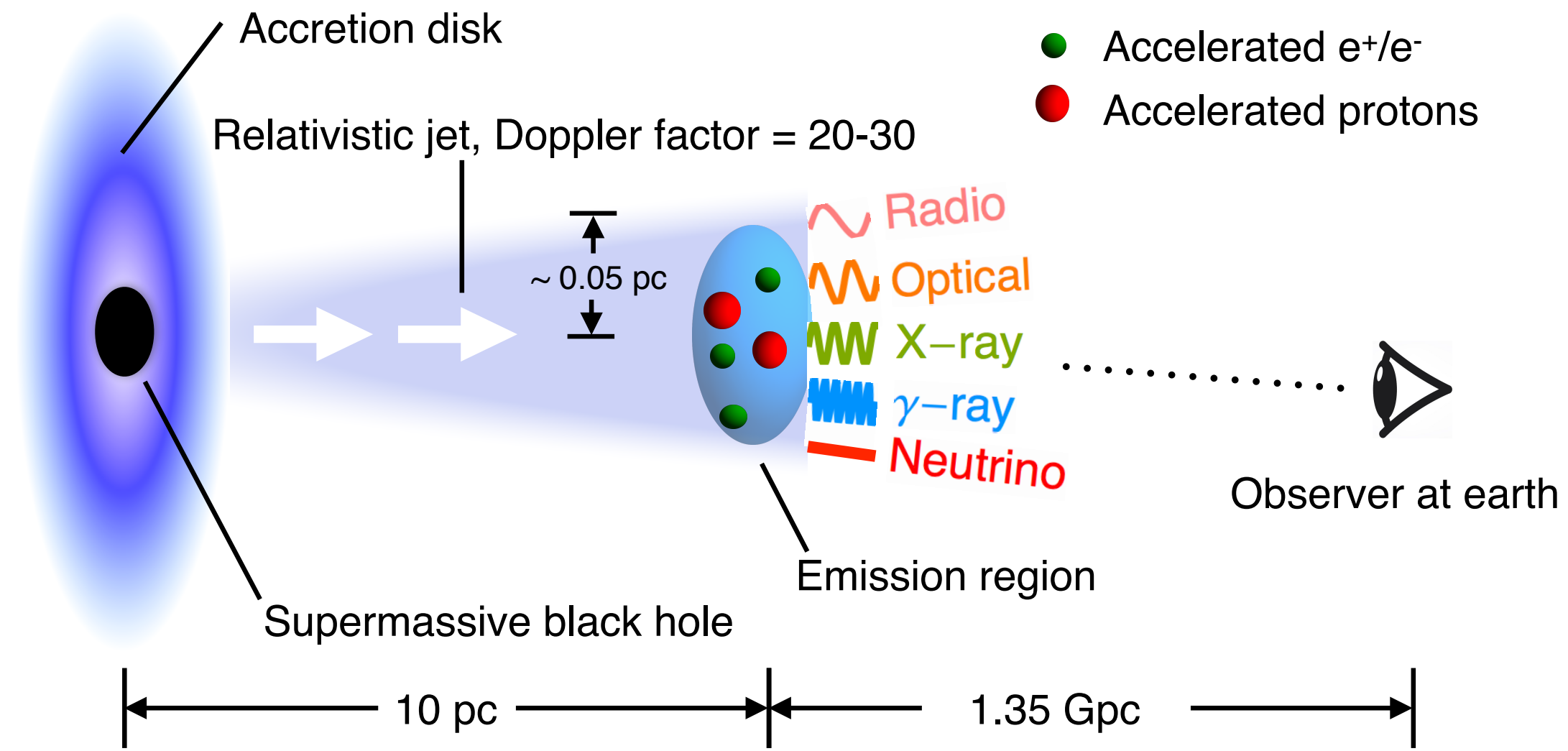
lab frame



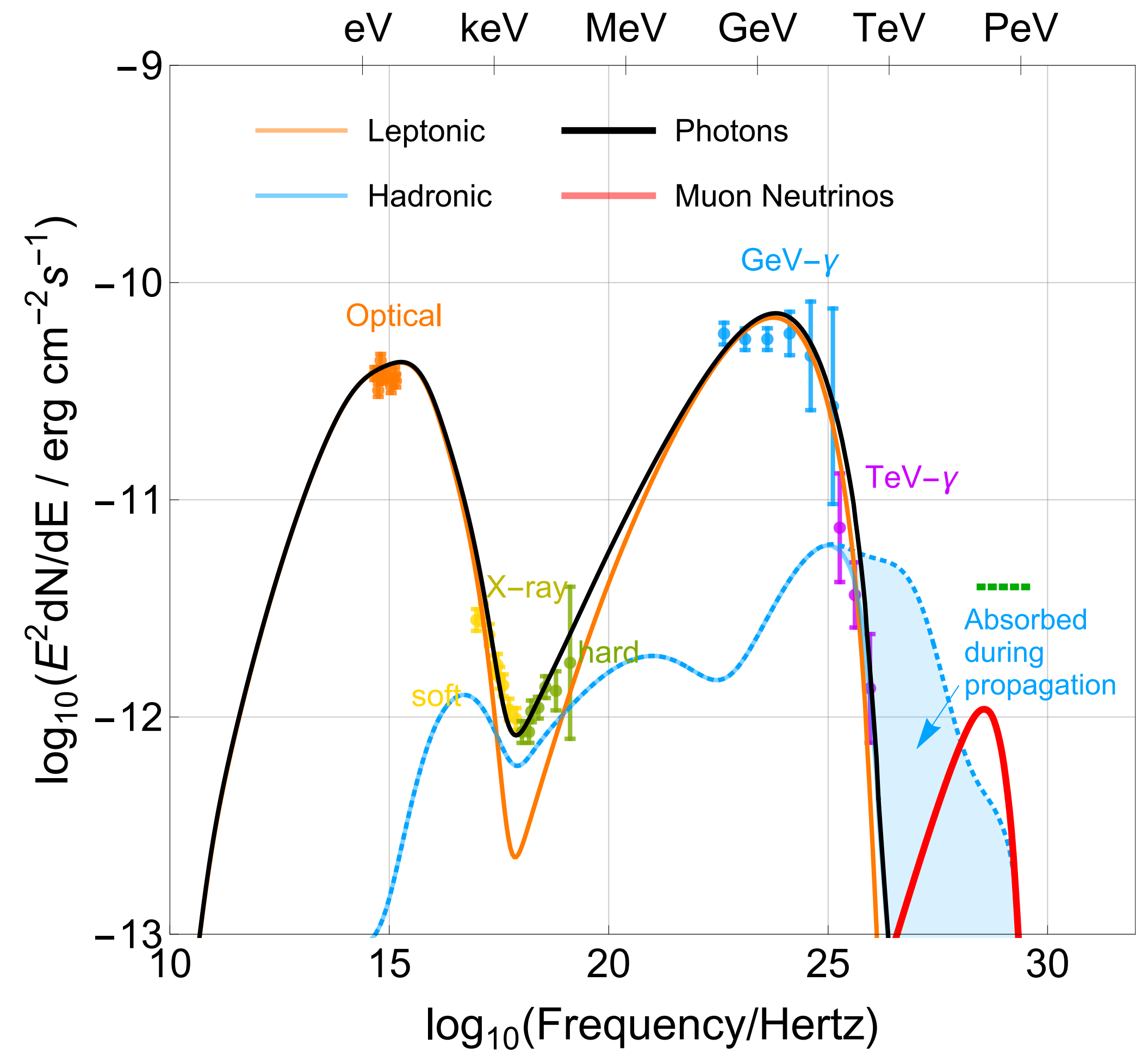
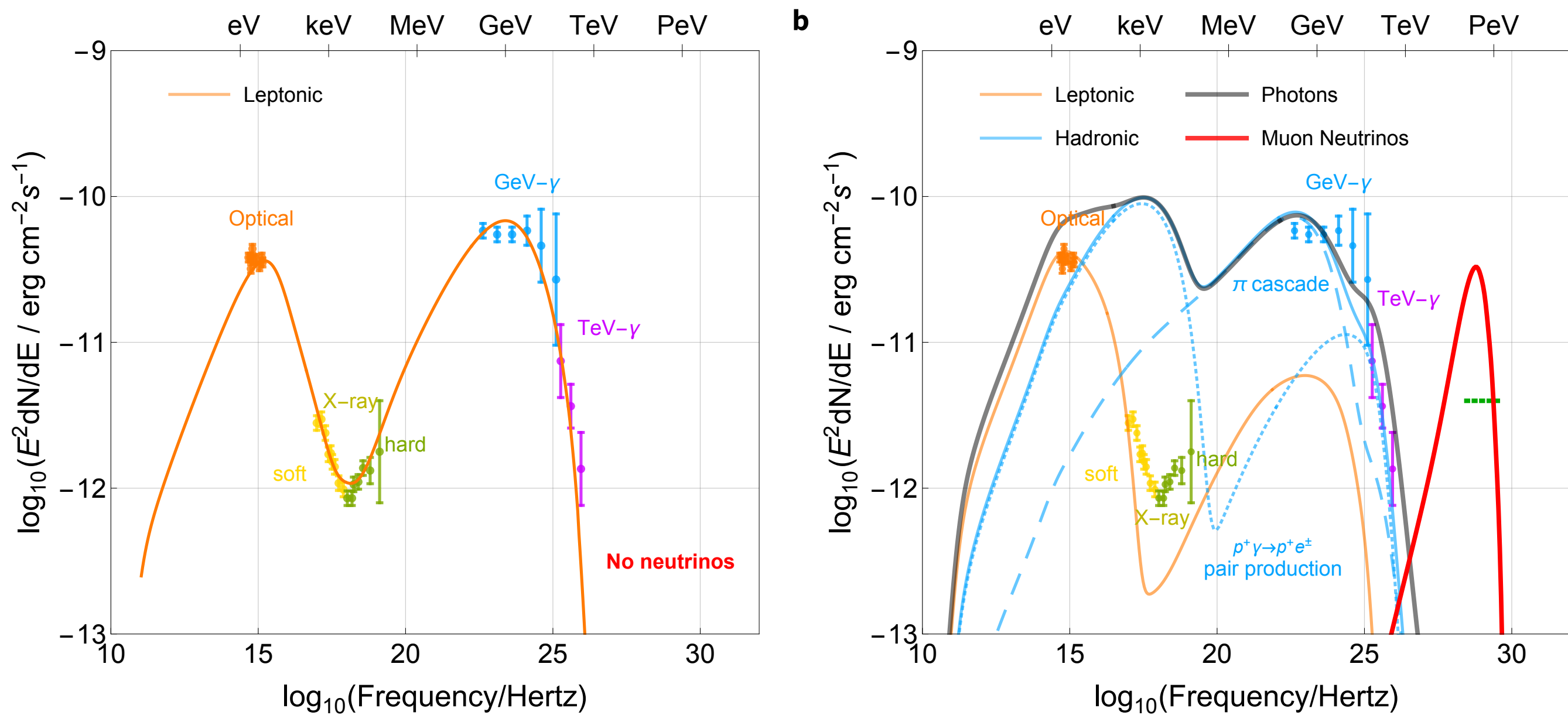
$$\frac{m_{\pi^0} c^2}{2} \sqrt{\frac{1-\beta}{1+\beta}} \leq E \leq \frac{m_{\pi^0} c^2}{2} \sqrt{\frac{1+\beta}{1-\beta}}$$

- ▶ cross section only mildly dependent on energy
- ◆ pions and gammas have roughly the same spectrum
- ▶ characteristic spectral feature observed in several sources

# leptonic vs. hadronic processes



- ▶ **blazar** TXS 0506+056
- ▶ correlated neutrino observation



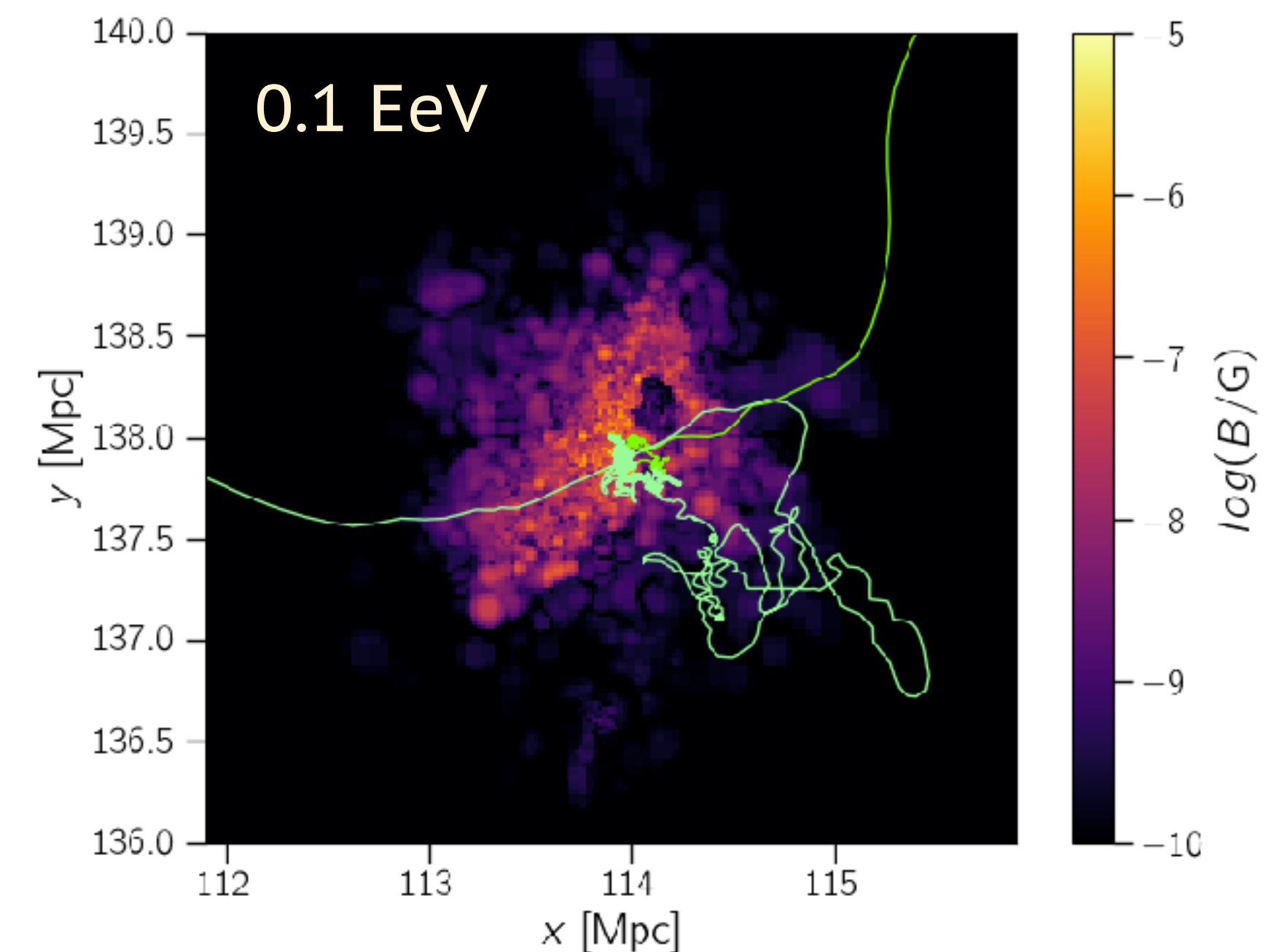
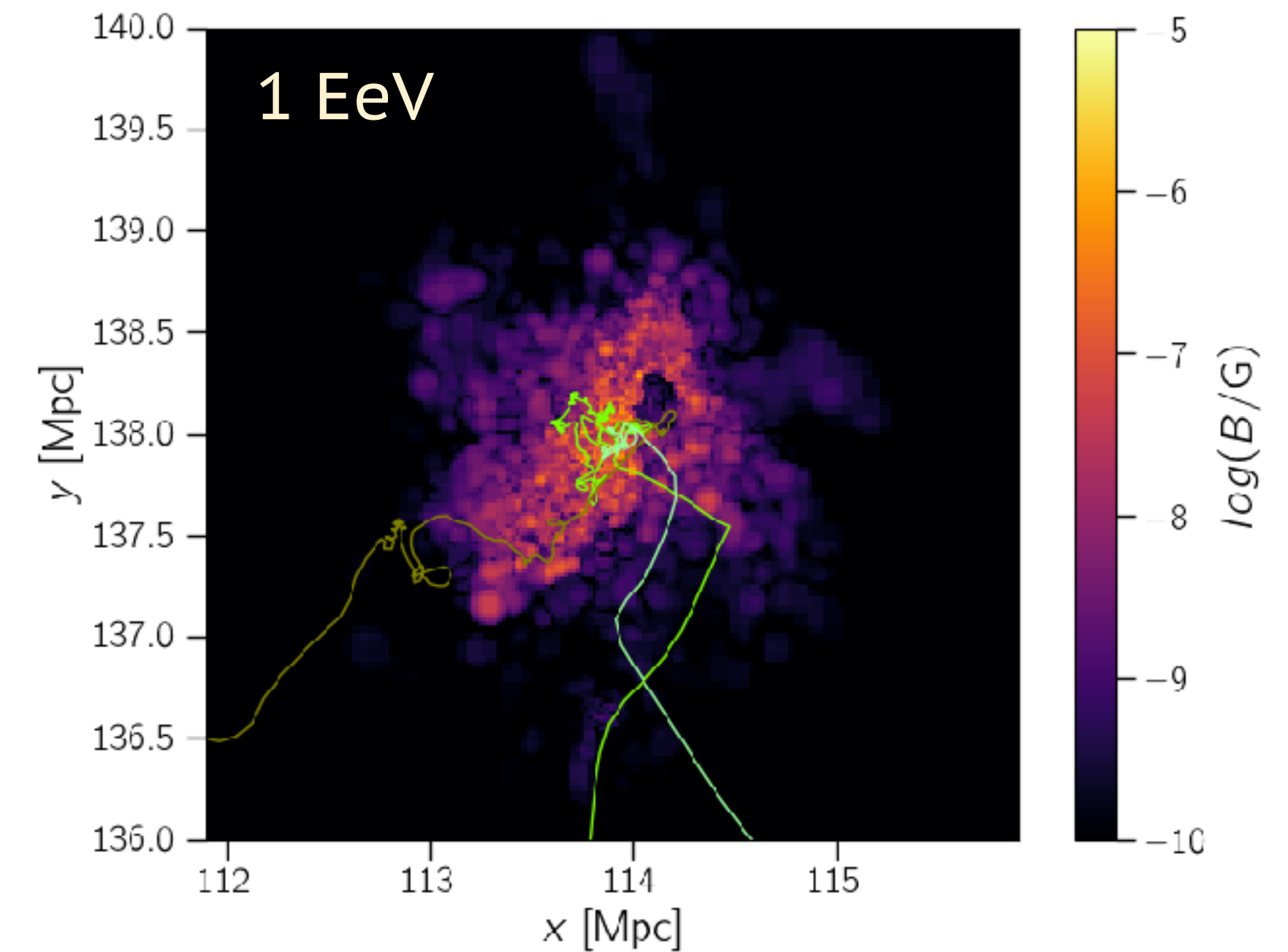
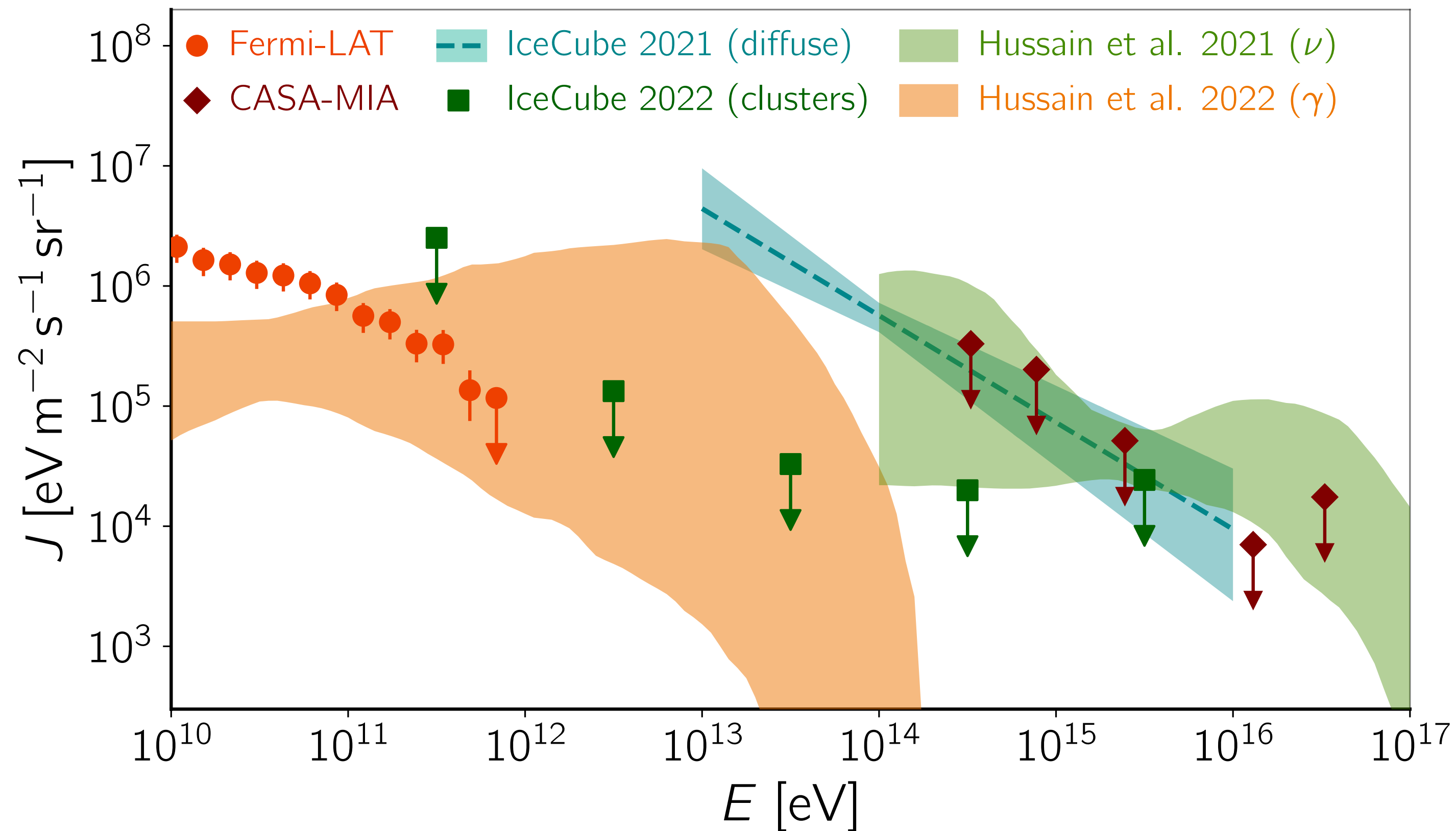


# multimessenger emission from galaxy clusters

Hussain, Alves Batista, de Gouveia Dal Pino, Dolag. MNRAS 507 (2021) 1762. arXiv:2101.07702

Hussain, Alves Batista, de Gouveia Dal Pino, Dolag. Nature Comms 14 (2023) 2486. arXiv:2203.01260

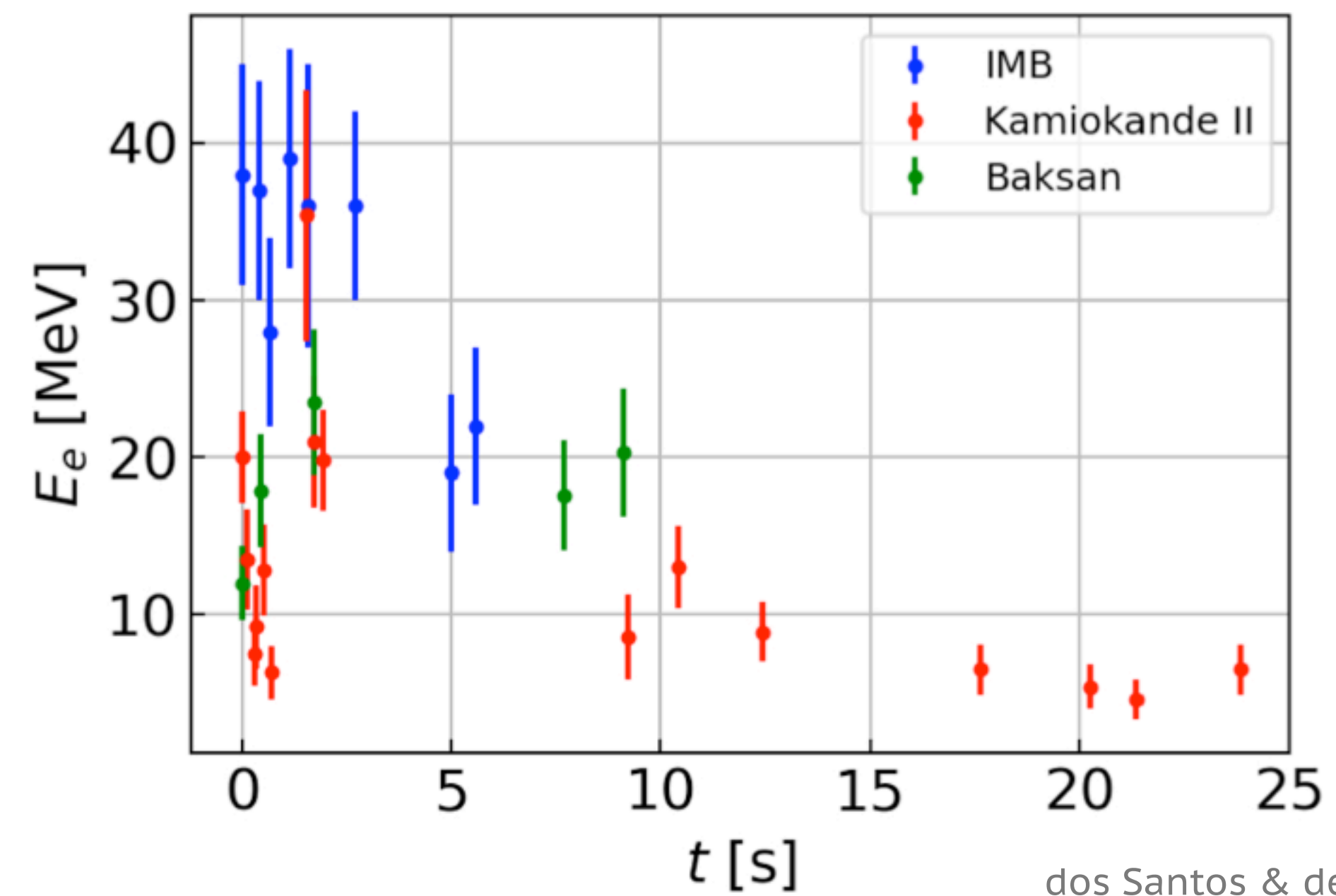
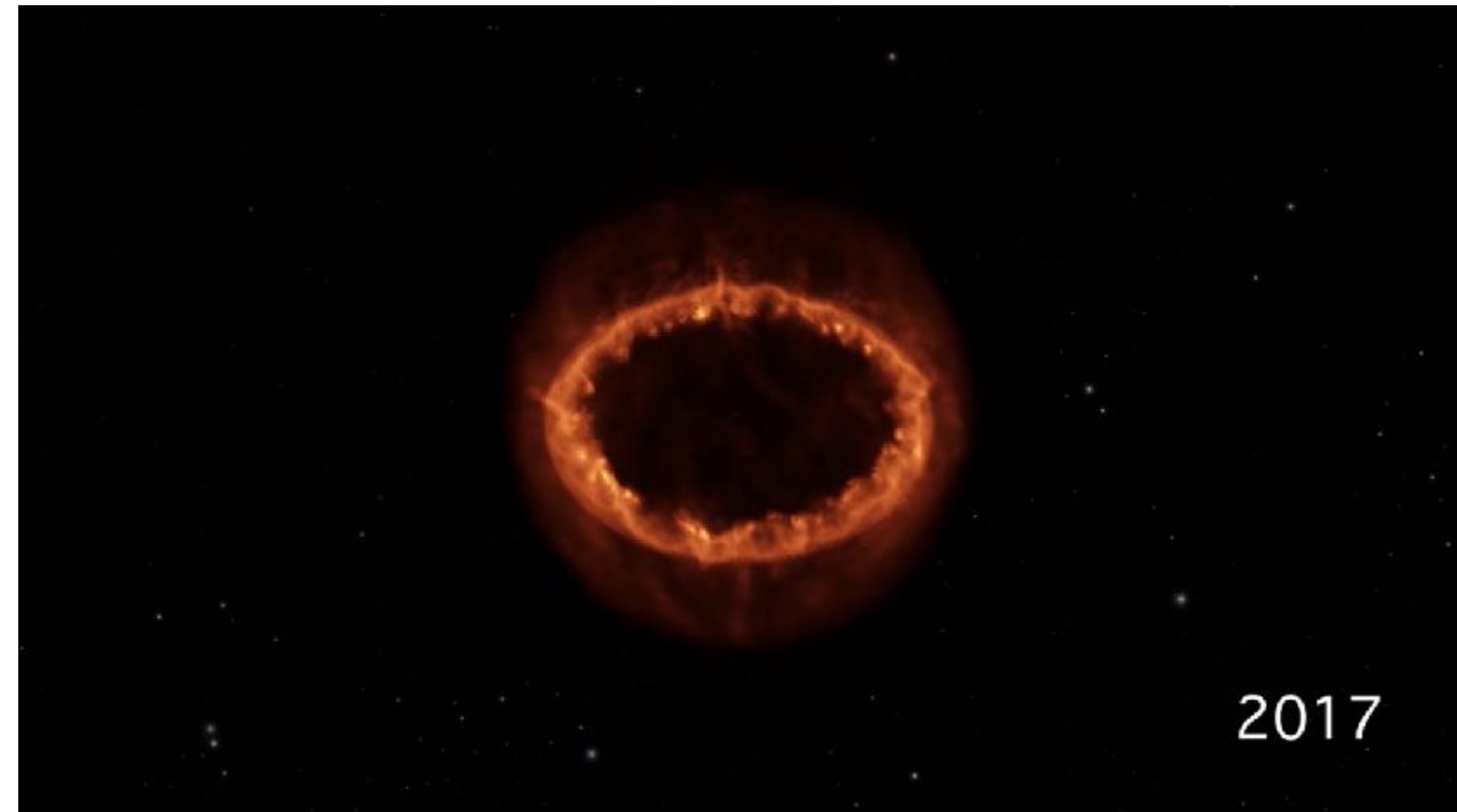
- ▶ high-energy neutrinos are produced via hadronic processes
- ▶ cosmic rays are affected by magnetic fields
- ▶ rate of CR interactions increases if magnetic fields are strong



# multi-messenger paradigm

# a few milestones. SN1987A

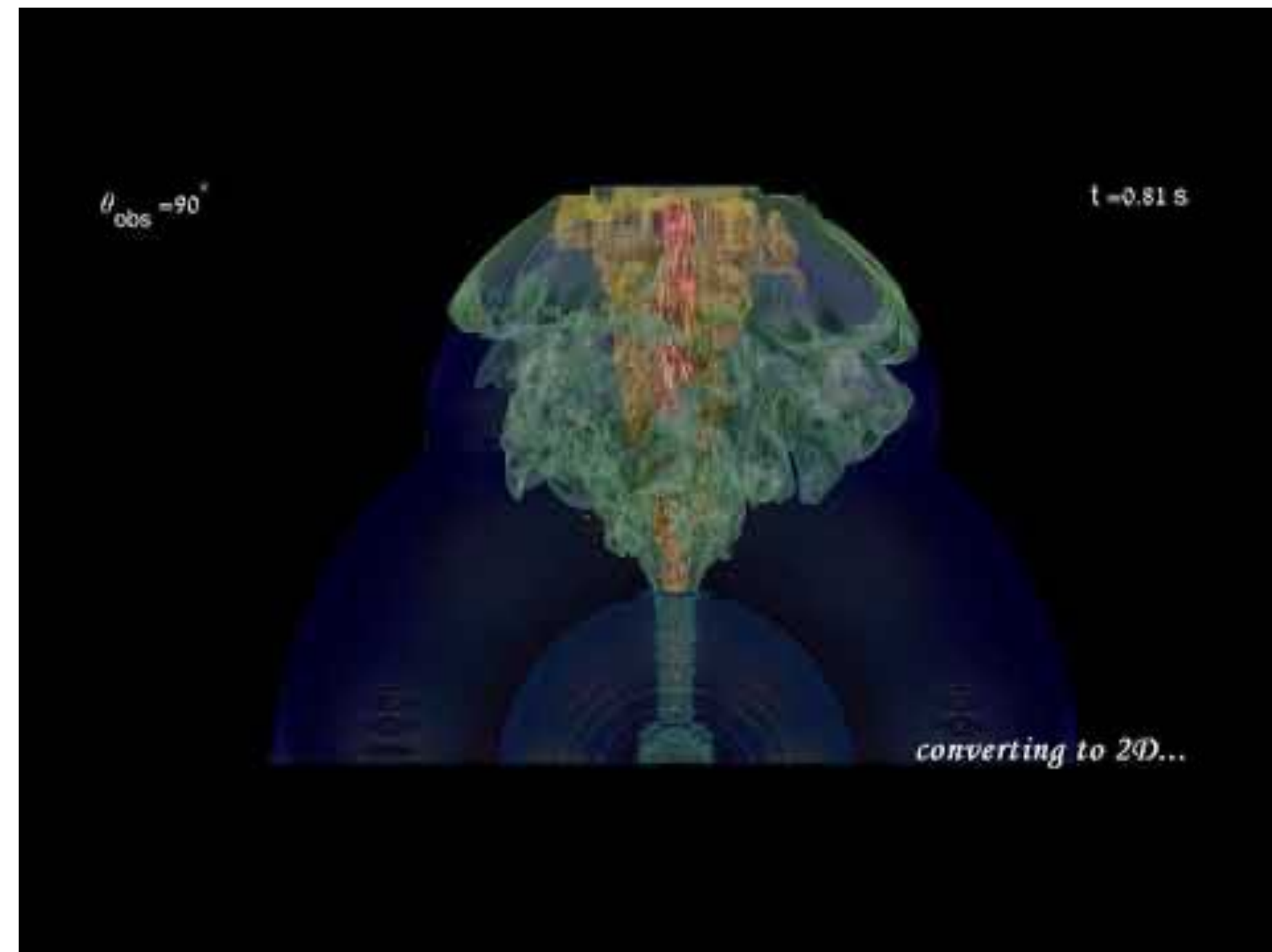
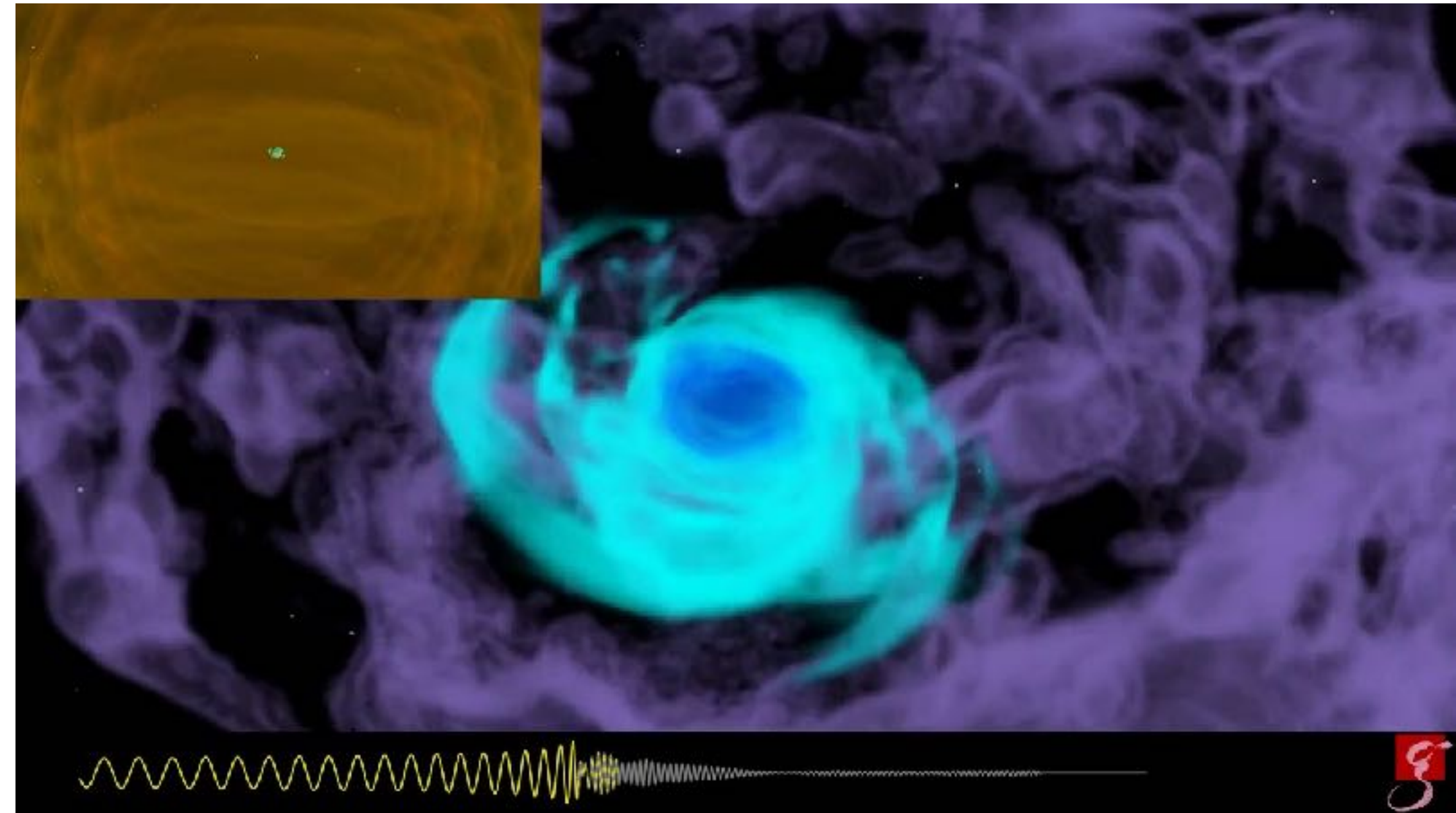
event	EM	CR	$\nu$	GW
SN 1987A	✓	x	✓	x
GW170817	✓	x	x	✓
TXS 0506+056	✓	x	✓	x
NGC 1068	✓	x	✓	x
Galactic plane	✓	✓	✓	x



# a few milestones. GW170817

event	EM	CR	$\nu$	GW
SN 1987A	✓	x	✓	x
GW170817	✓	x	x	✓
TXS 0506+056	✓	x	✓	x
NGC 1068	✓	x	✓	x
Galactic plane	✓	✓	✓	x

<https://youtube.com/watch?v=V6cm-0bwJ98>

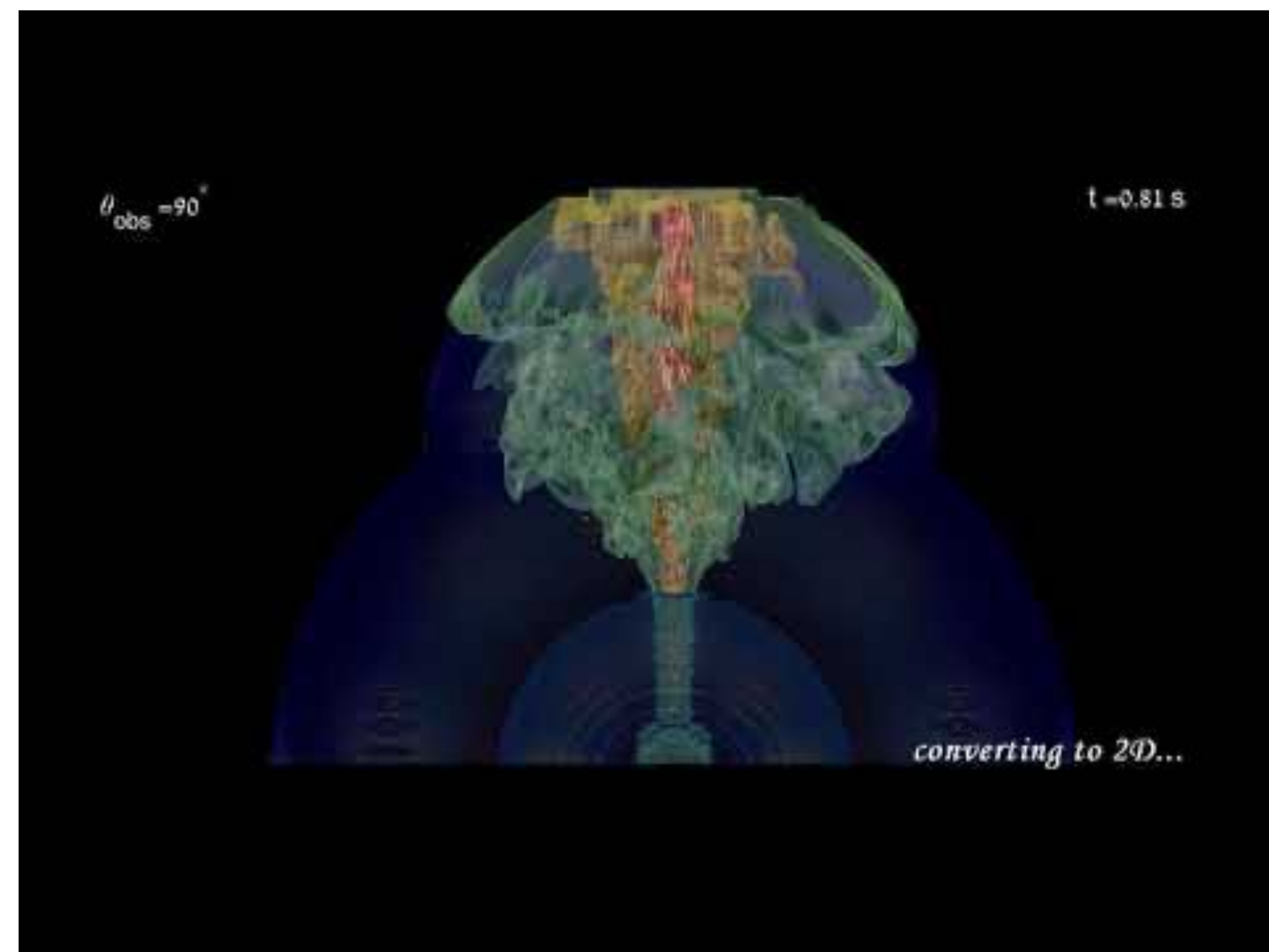
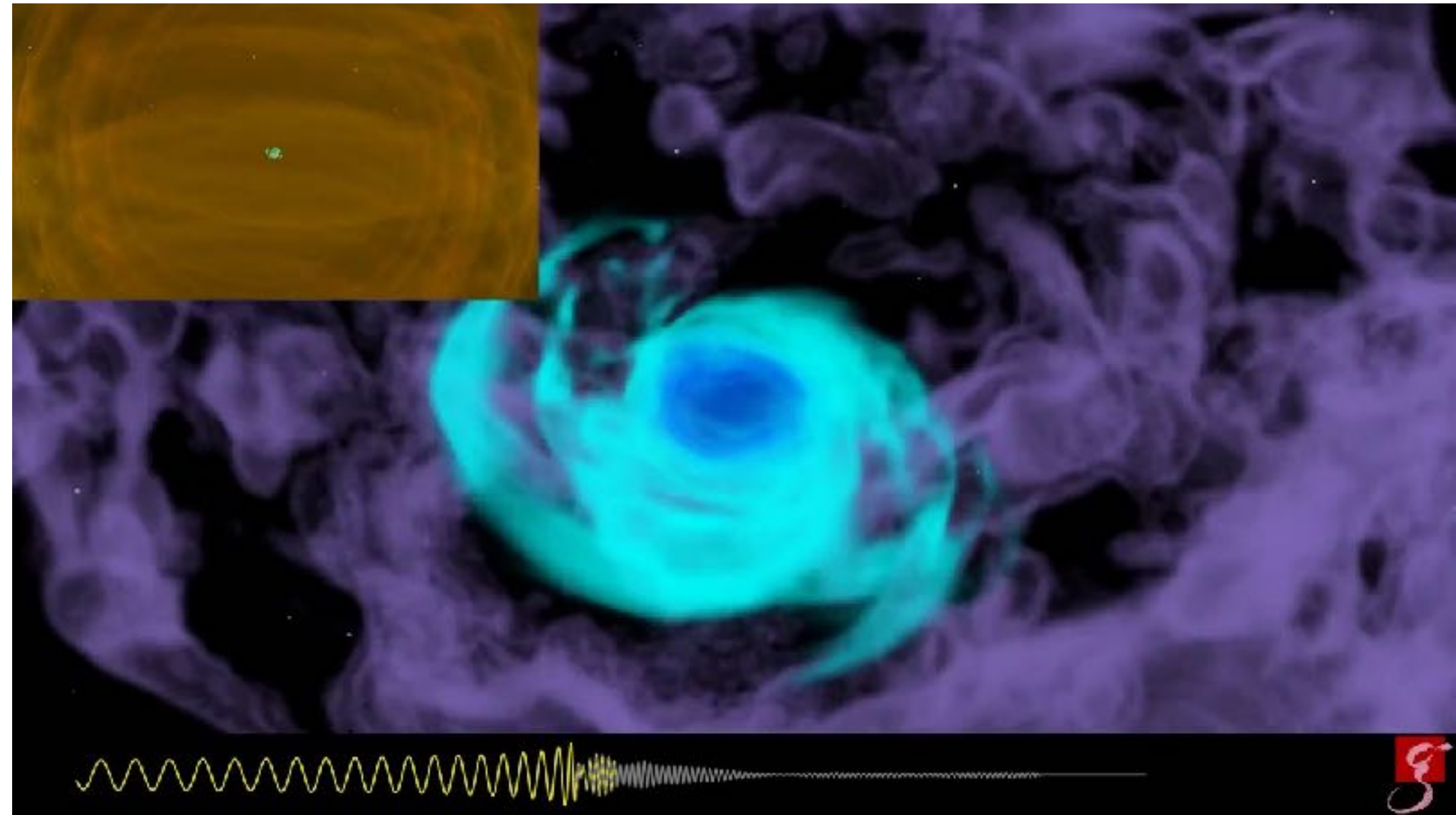


<https://www.youtube.com/watch?v=WoxmlY4Odlg>

# a few milestones. GW170817

event	EM	CR	$\nu$	GW
SN 1987A	✓	x	✓	x
GW170817	✓	x	x	✓
TXS 0506+056	✓	x	✓	x
NGC 1068	✓	x	✓	x
Galactic plane	✓	✓	✓	x

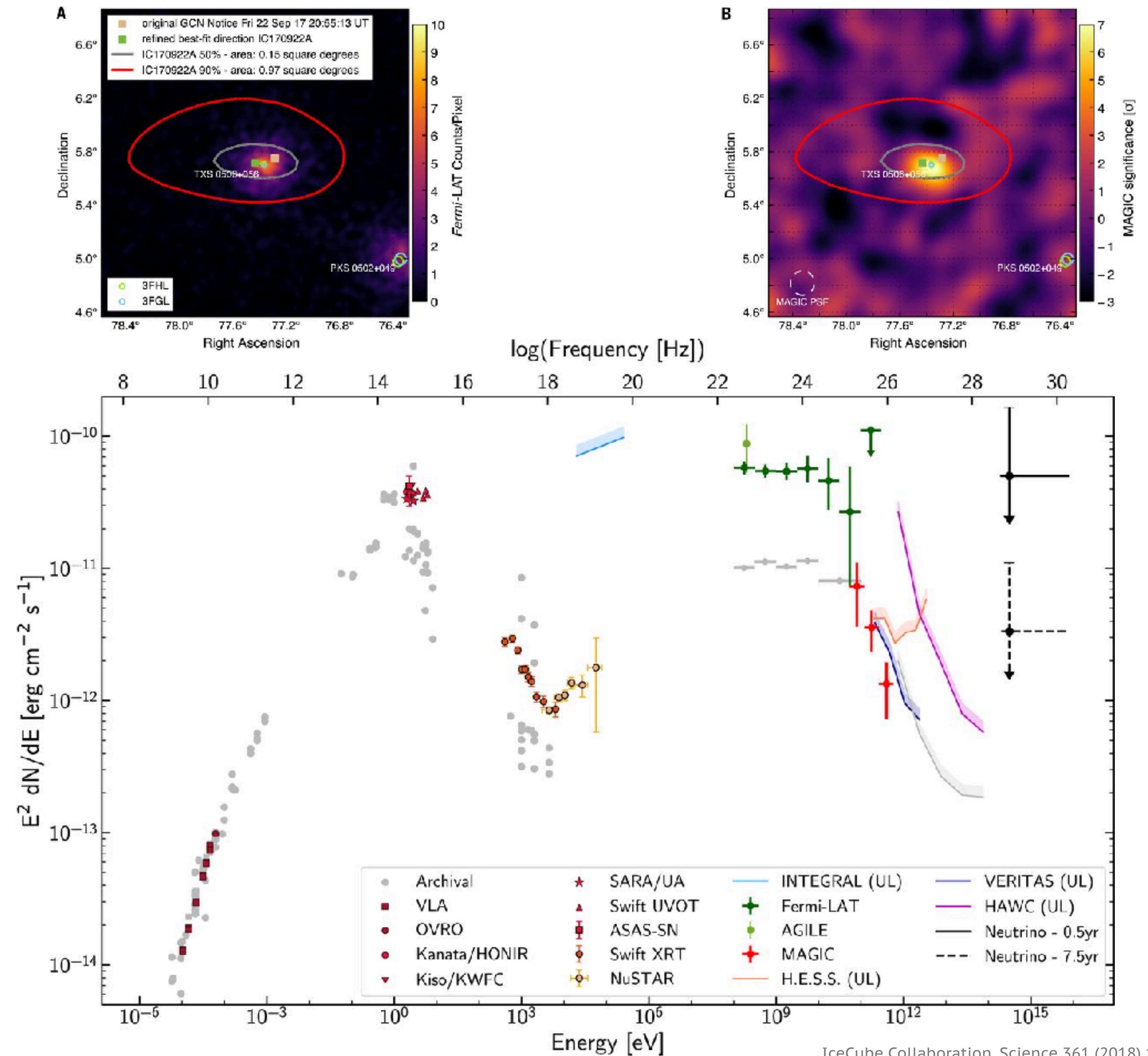
<https://youtube.com/watch?v=V6cm-0bwJ98>



<https://www.youtube.com/watch?v=WoxmlY4Odlg>

# a few milestones. TXS 0506+056

event	EM	CR	$\nu$	GW
SN 1987A	✓	X	✓	X
GW170817	✓	X	X	✓
TXS 0506+056	✓	X	✓	X
NGC 1068	✓	X	✓	X
Galactic plane	✓	✓	✓	X

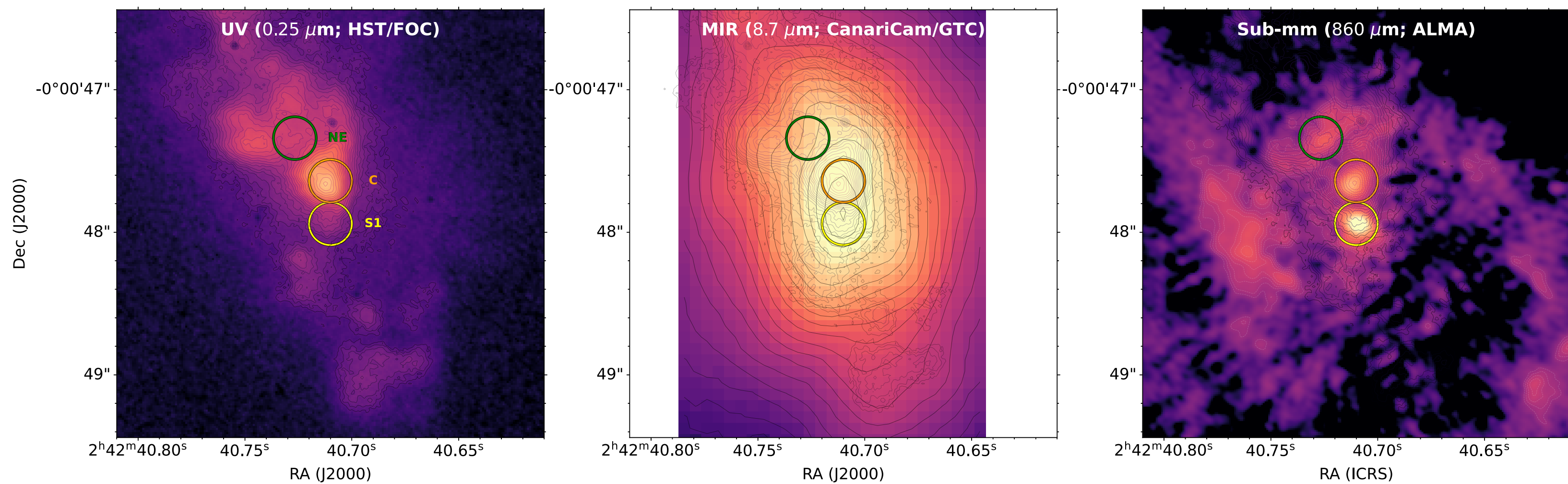
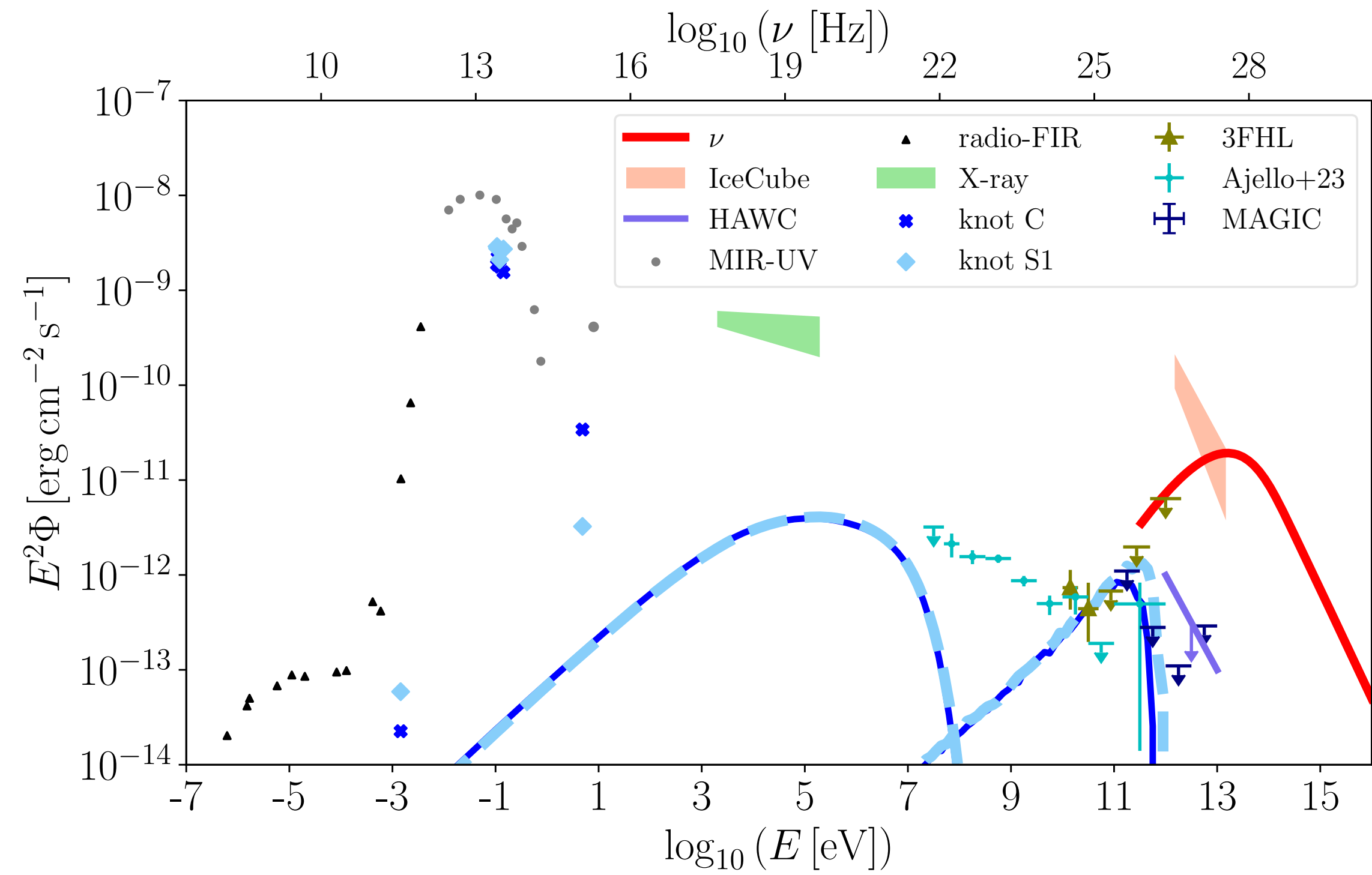


IceCube Collaboration. Science 361 (2018) 147. arXiv:1807.08794

IceCube Collaboration. Science 361 (2018) eaat1378. arXiv:arXiv:1807.08816

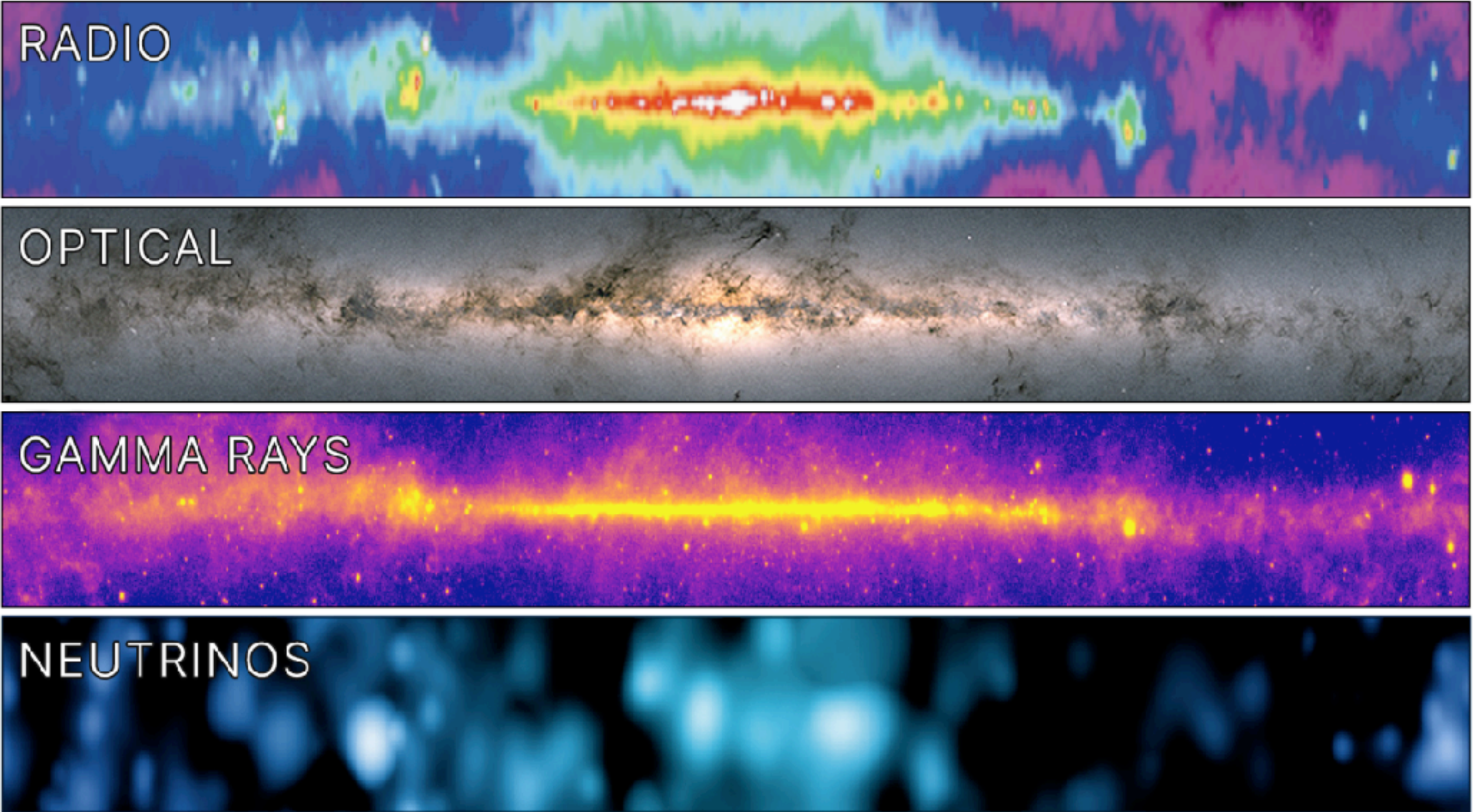
# a few milestones. NGC1068

event	EM	CR	$\nu$	GW
SN 1987A	✓	✗	✓	✗
GW170817	✓	✗	✗	✓
TXS 0506+056	✓	✗	✓	✗
NGC 1068	✓	✗	✓	✗
Galactic plane	✓	✓	✓	✗



# a few milestones. Galactic plane

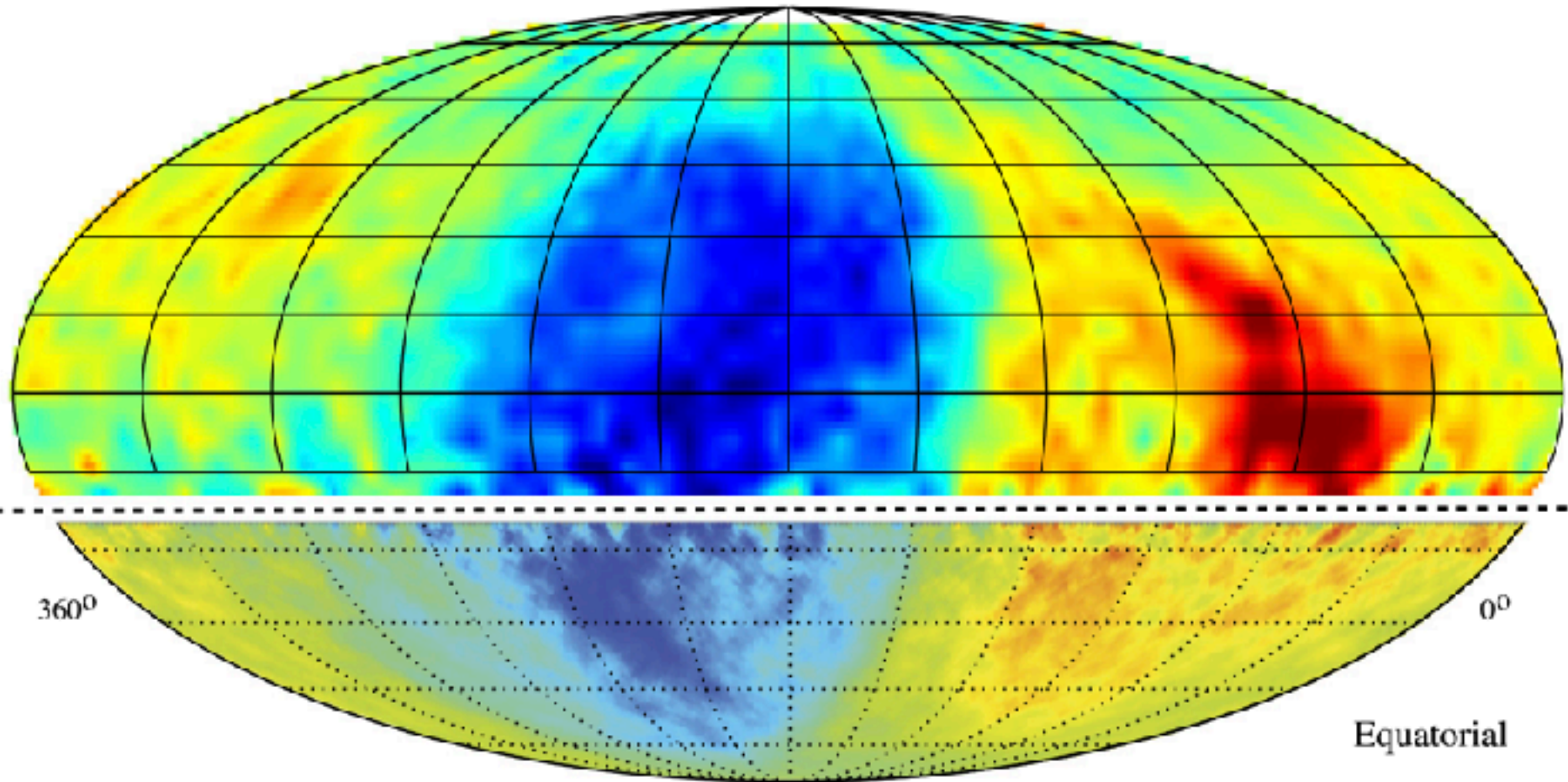
event	EM	CR	$\nu$	GW
SN 1987A	✓	x	✓	x
GW170817	✓	x	x	✓
TXS 0506+056	✓	x	✓	x
NGC 1068	✓	x	✓	x
Galactic plane	✓	✓	✓	x



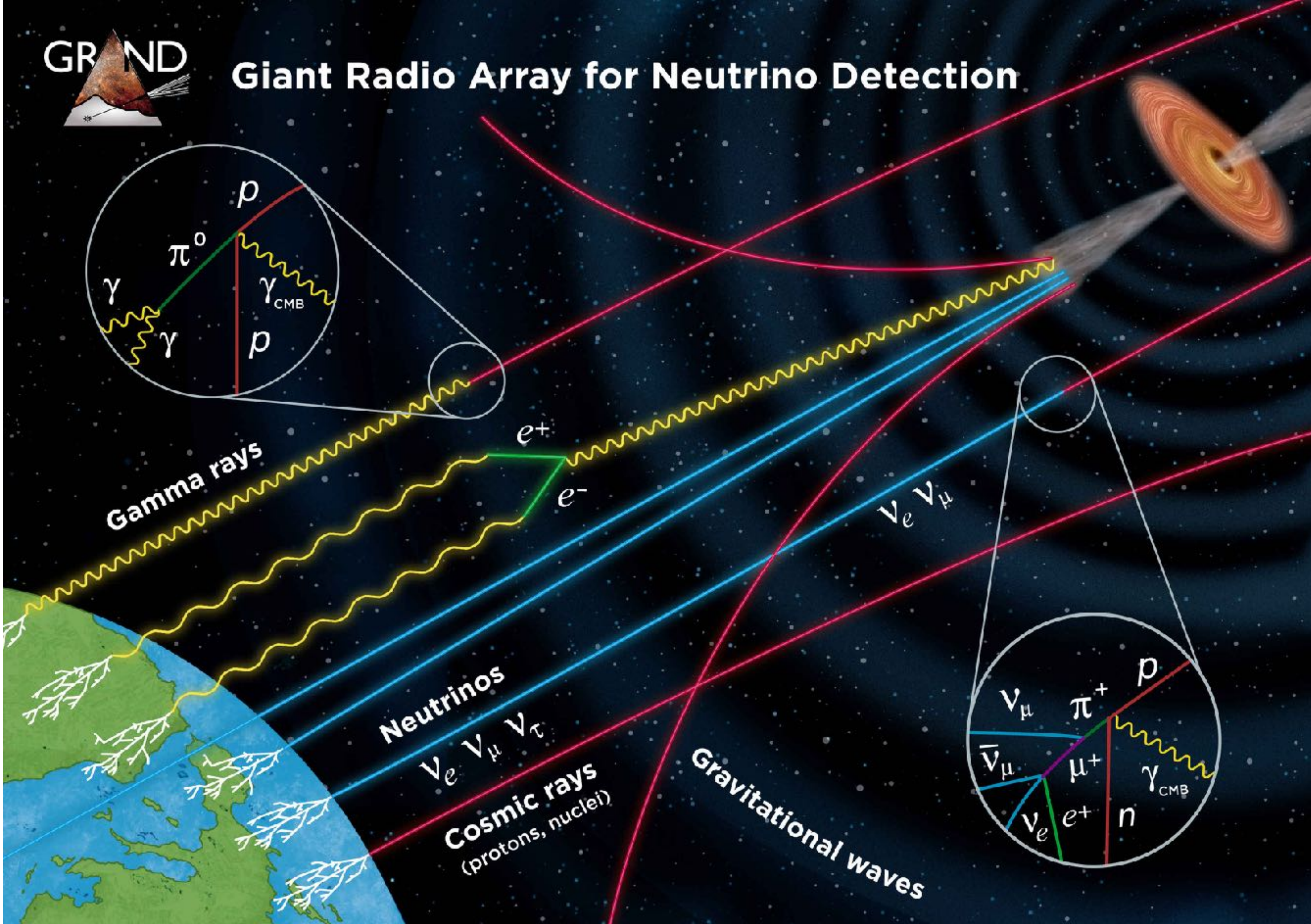
## CRs

Tibet  
5 TeV

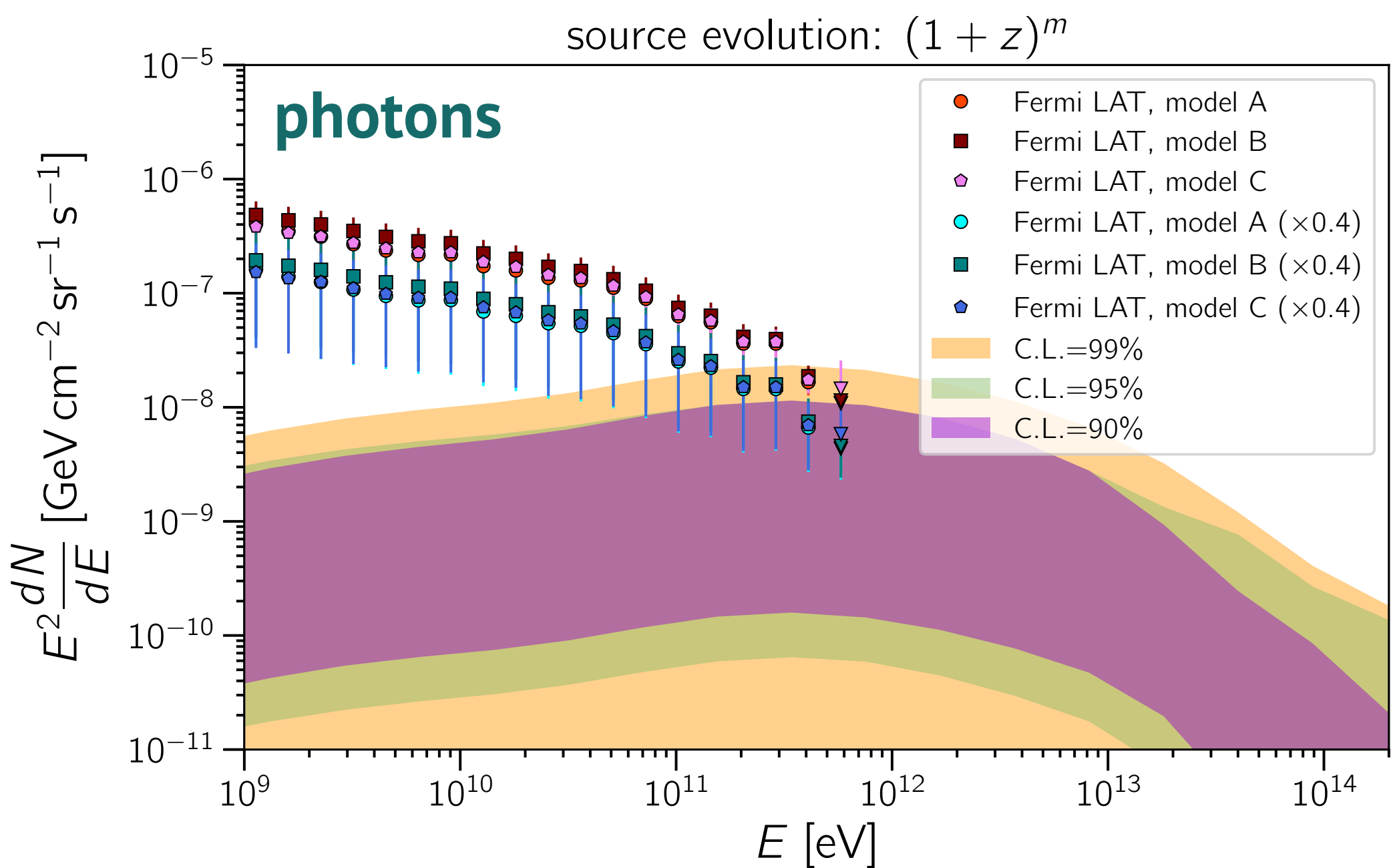
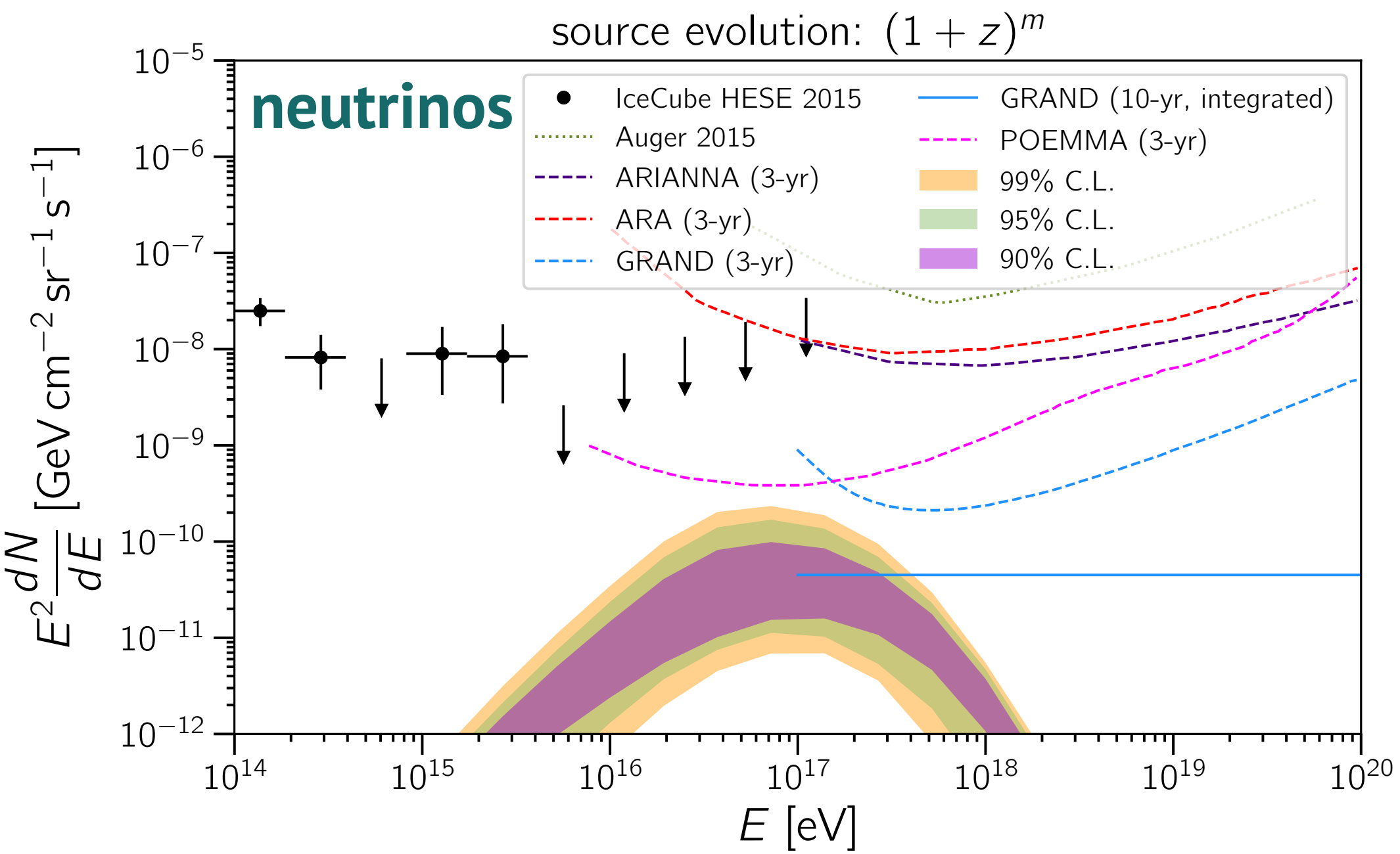
IceCube  
13 TeV



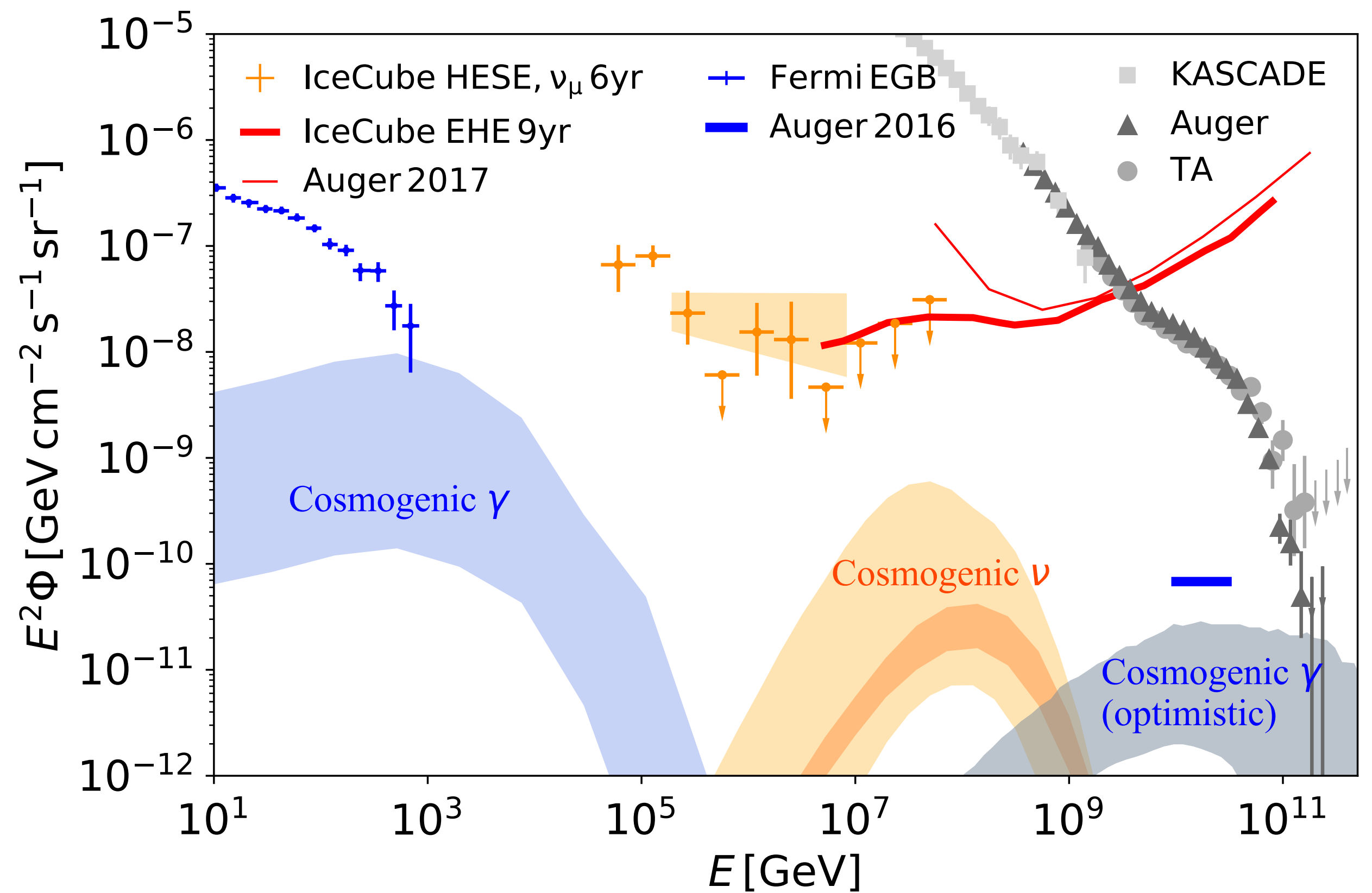




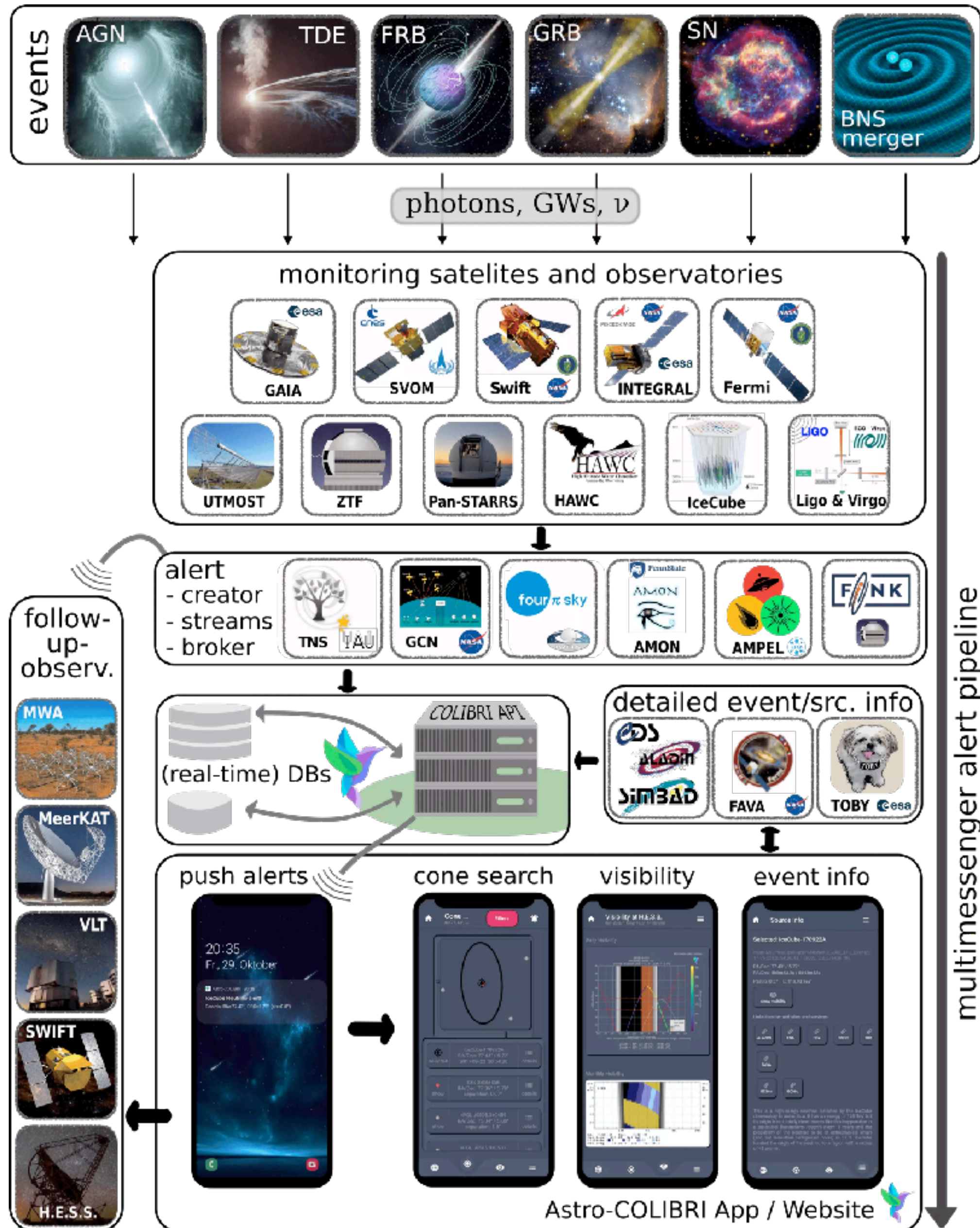
# cosmogenic particles



- ▶ no cosmogenic neutrinos have been confirmed yet
- ▶ only one neutrino with  $E > 100$  PeV
- ▶ no confirmed detection of cosmogenic photons



# multi-messenger networks and alert systems



Astro-COLIBRI

